

1 Estimation of surface iron oxide abundance with
2 suppression of grain size and topography effects

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22

23 **Abstract**

24 Mineral forms of iron oxide, such as hematite, goethite and jarosite, are important
25 because they are widely distributed at the Earth's surface and because they are used as
26 indicators for mineral exploration. Iron oxide abundance in rocks containing these
27 minerals can be estimated from the absorption depth at wavelengths of around 900 nm in
28 a reflectance spectrum, but this depth is also affected by extraneous factors such as grain
29 size and topography. This paper investigated the effect of grain size on reflectance spectra
30 and proposed a method for estimating iron oxide abundance in surface rocks by using
31 remotely sensed data with suppression of the effects of grain size and topography.
32 Reflectance spectra were measured in a laboratory from rock powder samples of different
33 grain sizes containing iron oxide minerals. While the reflectance increased with
34 decreasing grain size, the presence of ferric iron caused the absorption depth to be almost
35 constant at around 900 nm, irrespective of the chemical composition of the sample. In
36 addition, the difference between the reflectance at 550 nm and 760 nm (Slope) was a
37 function of grain size. Iron oxide abundance can be estimated accurately by MCR-900D,
38 which is the maximum absorption depth at the absorption center after the effect of grain
39 size and topography was suppressed by Slope and the continuum-removal method, which
40 takes the ratio between the original spectrum and its continuum, respectively. Correlation
41 of MCR-900D results with datasets of actual spectral and chemical iron oxide laboratory
42 measurements revealed that the mineral forms also need to be considered. MCR-900D
43 results were significantly correlated with rock samples classified as containing different
44 forms of iron oxide minerals (hematite, goethite and jarosite). Finally, MCR-900D was
45 applied to an AVIRIS dataset for the Cuprite site in Nevada, USA. The results represented
46 the enrichment zones of iron oxide within hydrothermally altered areas.

47

48 Key words: iron oxide; reflectance spectra; grain size; remote sensing; AVIRIS

49

50 **1. Introduction**

51 Mineral forms of iron oxide, such as hematite, goethite and jarosite, are widely
52 distributed at the Earth's surface. The presence of ferric iron in these minerals means that
53 they have three diagnostic absorption features in the visible to near-infrared (VNIR)
54 regions; they exhibit charge transfer absorption between 480 and 550 nm, and crystal-
55 field absorption between 630 and 715 nm, and between 850 and 1000 nm (Morris et al.,
56 1985). These spectral features can be used to identify and distinguish between iron oxide
57 minerals. This study considered iron oxide minerals to include hematite (α -Fe₂O₃),
58 goethite (α -FeOOH) and jarosite (KFe₃(SO₄)₂(OH)₆). In addition, iron oxide (i.e., ferric
59 oxide or Fe₂O₃) abundance was considered to include iron oxyhydroxide (FeO(OH))
60 abundance, which is typically measured as Fe₂O₃ in chemical analysis. Most of the
61 studies conducted on these spectral reflectance characteristics to date have been
62 performed in laboratories (Morris et al., 1985; Sherman and Waite, 1985). Studies
63 conducted in the field have shown that iron oxide minerals are not randomly distributed,
64 but that they occur in zones associated with ore deposits (e.g., around ore bodies of
65 porphyry copper deposits) (Townsend, 1987). Consequently, if the distribution of iron
66 oxide minerals is associated with certain ore minerals (Gustafson and Hunt, 1975; Saegart
67 et al., 1974), then locating minerals containing iron oxide can be used to identify sites
68 that may be suitable for mineral exploration (Ciampalini et al., 2013).

69 In the field of planetary science, remote sensing data has been used to estimate FeO
70 abundance. Iron exists as ferrous oxide on the surface of the moon and as ferric oxide on

71 the Earth. Lucey et al. (1998) estimated the FeO content of areas on the lunar surface after
72 suppressing space weathering effects. On the other hand, the quantitative estimation of
73 ferric iron oxide abundance on the Earth's surface by remote sensing is still in the initial
74 stages. Fe₂O₃ is the most common iron oxide, and its abundance affects the depth for
75 absorption features between 850 and 1000 nm (Cudahy and Ramanaidou, 1992; Hunt and
76 Ashley, 1979). However, since the spectral reflectance of rocks and minerals in remote
77 sensing data is also affected by factors such as chemical composition, topography, mixing
78 and grain size (Murphy and Monteiro, 2013), these factors can affect the absorption depth,
79 which is defined as the difference between the minimum reflectance of an absorption and
80 a reflectance at the same wavelength on the straight line connecting two shoulders on the
81 both sides of an absorption. Given the complexity of the interactions among these factors,
82 their relative contribution to estimates of iron oxide abundance is typically not
83 considered; however, we consider they would amount to several wt%.

84 Previous studies that have attempted to infer iron oxide abundance based on the
85 reflectance spectra of surface rocks and/or rock samples have typically employed band
86 ratios to suppress the effect of topography. For example, Advanced Spaceborne Thermal
87 Emission and Reflection Radiometer (ASTER) band4/band3 is used for estimating iron
88 oxide abundance, and ASTER band2/band1 is used for classifying iron oxide forms
89 (Cudahy, 2012). The band ratio is effective for removing the effect of topography;
90 however, when the slope of the reflectance values of two spectral bands changes due to
91 other effects, such as water content, organic material or grain size, using band ratios is
92 not effective for removing the effect of topography (Murphy and Monteiro, 2013).
93 Although Haest et al. (2012) managed to suppress the contribution of brightness by using
94 reflectance at 1650 nm, which was assumed to be a function of brightness, reflectance at

95 that wavelength would also be affected by the mineral composition and topography. The
96 continuum-removal method (Clark and Roush, 1984) is a powerful pre-processing
97 procedure that enhances absorptions in reflectance spectra in order to accurately identify
98 minerals, rocks and vegetation (e.g., Gomez et al., 2008; Oshigami et al., 2013; Schmidt
99 et al., 2003). In addition, the method is well suited to identifying minerals and rocks,
100 particularly since the processed spectra are not affected by topography. However, the
101 absorption depth for minerals identified using data processed with the continuum-removal
102 method increases with increasing grain size (Clark, 1999). There are thus no established
103 methods for estimating iron oxide abundance in remote sensing data that can suppress the
104 contribution of grain size and topography.

105 The primary objective of this study was therefore to develop a method for estimating
106 iron oxide abundance in surface rocks while suppressing the effects of grain size and
107 topography. Accurate estimates of iron oxide abundance are necessary, not only to
108 characterize the geology of an area of interest, but also because iron oxide abundance is
109 used as an indicator of various ore deposits in exploration geology.

110 Reflectance spectra of samples containing iron oxides were collected, and the relation
111 between grain size and absorption depths at around 900 nm due to iron oxides were
112 investigated in the laboratory. The absorption depth was corrected by the continuum-
113 removal method and the slope of the reflectance spectrum from 550 to 760 nm was plotted
114 to suppress the topography and grain size effect, respectively. The relations between the
115 absorption depths and iron oxide abundance were then established for three iron oxide
116 minerals, and these relations were then applied to actual airborne image data of Cuprite,
117 Nevada, USA obtained by AVIRIS. We discussed the effectiveness of our method for
118 suppressing the grain size and topography effect when estimating iron oxide abundance.

119

120 **2. Approach**

121 *2.1. Absorption depth*

122 The absorption depth for rocks and minerals in this study was obtained by subtracting
123 the continuum of a spectrum from the original spectrum to obtain a normalized absorption
124 depth. The continuum is a convex background that decreases toward the shorter and
125 longer wavelength regions, and can be drawn as a straight line on the reflectance spectra
126 of iron oxide minerals in the region from 760 nm to 1250 nm. Thus, 900D, which is the
127 maximum absorption depth between 850 and 1000 nm, can be calculated using the
128 following equation:

129

$$130 \quad 900D = 900C - 900R \quad (1)$$

131

132 where 900C is a continuum value of a spectrum at approximately 900 nm, and 900R is
133 the original spectral reflectance at around 900 nm. We assumed that the continuum
134 represents the slopes of absorption values centered outside the 900 nm region due to other
135 chemical components. In this way, the continuum-removal method would allow us to
136 examine absorption by iron oxide only centered at around 900 nm.

137 In order to suppress the effects of grain size and topography, we introduced two
138 absorption depths modified from 900D. CR-900D stands for Continuum-Removed 900D
139 and indicates an absorption depth obtained from the reflectance spectrum after
140 suppression of the topography effect by the continuum-removal method. MCR-900D
141 stands for Modified CR-900D and indicates an absorption depth obtained from the
142 reflectance spectrum after suppression of the effects of grain size and topography by the

143 slope and the continuum-removal methods, respectively. The methods for calculating CR-
144 900D and MCR-900D are explained in the following sections.

145

146 2.2. Topographic correction

147 At wavelengths longer than 550 nm, spectral reflectance increases at a constant rate as
148 grain size decreases, but 900D does not change. In other words, differences in the
149 reflectance values of samples with two different grain sizes are constant in this
150 wavelength region. On the other hand, topography changes both spectral reflectance and
151 900D. The spectral reflectance of minerals and rocks is multiplicatively affected by
152 topographic slopes (Green and Craig, 1985). If a the topographic slope is assumed to be
153 $\tilde{\alpha}$, then the topography effect is indicated as $T(\tilde{\alpha})$ (e.g., $T(\tilde{\alpha}) = \cos^2\tilde{\alpha}$). The apparent
154 reflectance with the topography effect (R_{topo}) can be expressed by the following
155 equation:

156

$$157 \quad R_{topo} = R \times T(\tilde{\alpha}) \quad (2)$$

158

159 where R is the original spectral reflectance, and $T(\tilde{\alpha})$ affects the spectral reflectance in all
160 wavelength regions at a constant ratio regardless of the material in a pixel.

161 The continuum-removal method of Clark and Roush (1984) can then be applied to
162 suppress the effect of topography. A continuum-removed spectrum can be derived by
163 dividing an original spectrum (R) by its continuum (C), and the absorption depth due to
164 ferric iron can generally be obtained by the following equation:

165

$$166 \quad CR-900D = 1.0 - 900R / 900C \quad (3)$$

167

168 where CR-900D is the absorption depth at an absorption center around 900 nm formed
169 by ferric iron in a continuum-removed spectrum, 900R is the original spectral reflectance
170 at the absorption center of the original spectrum, and 900C is the continuum value of the
171 original spectrum at the absorption center in the wavelength region of 760 nm to 1250
172 nm. The CR-900D value of 0.0 means that the original reflectance matches its continuum
173 (i.e. the reflectance peak between two absorptive or non-absorptive regions). The CR-
174 900D is not affected by the effect of topography, because both the original spectrum and
175 its continuum change at the same rate $T(\lambda)$ in equation (2), thus $900R/900C$ does not
176 change either by the effect of topography.

177

178 *2.3. Reflectance measurements for defining the grain size effect*

179 Five rock (A, B, E, F, G in Table 1 and Figure 1) and two sand (Sample C, D) samples
180 containing iron oxide minerals were used to investigate spectral changes attributed to
181 different grain sizes. These samples were collected in Utah and Nevada, USA.; four
182 contained hematite, two contained goethite, and one contained jarosite. The sand samples
183 were divided by sieving, and the remaining rock samples were powdered using a ball mill
184 followed by sieving. The sand samples were not powdered, because they consisted of fine
185 quartz grains which were ready for sieving. Grain-size bins for sieving were <25 μm , 25-
186 44 μm , 44-75 μm , 75-125 μm , 125-180 μm , 180-250 μm , and 250-500 μm . Some grain
187 size fractions could not be prepared because of low sample volumes.

188 Reflectance spectra of each sample with different grain sizes were measured in the
189 laboratory using a spectroradiometer (FieldSpec Pro, ASD Inc., Co) which has 2150
190 contiguous spectral bands with wavelengths ranging from 350 to 2500 nm, and a spectral

191 resolution of 3 nm at 700 nm and 30 nm at 1400 and 2100 nm. In this study, reflectance
192 in laboratory measurements refers to the ratio between radiance from a target and radiance
193 from a Lambertian surface under the same measurement conditions. A Spectralon panel
194 (Labsphere Inc., NH) was used as a reference and was assumed to be a Lambertian surface.
195 A halogen lamp was used as an illumination source, but reflectance spectra in the shorter-
196 wavelength region were noisy due to the low light intensity. Therefore, the spectral data
197 for wavelengths shorter than 450 nm were not used in this study. Reflectance spectra of
198 each sample were measured five times, and the reflectance and absorption depth values
199 were averaged. The RMS error of five measurements for each sample was less than 1%
200 of reflectance at wavelengths longer than 450 nm. The chemical compositions of all
201 samples were also analyzed by X-Ray fluorescence (XRF) spectroscopy (ZSK Primus II
202 instrument equipped with a Rh mode, Rigaku, Japan) (Table 1).

203 Figure 1 shows reflectance spectra obtained for samples of different grain sizes.
204 Changes in reflectance spectra with grain size differ according to whether the minerals
205 being analyzed are transparent, opaque or trans-opaque (Salisbury and Hunt, 1968). As
206 the grain size decreases, the reflectance of transparent minerals (e.g., silicates) increases,
207 while that of opaque minerals (e.g., metal sulfides) decreases. Some iron oxide minerals,
208 such as goethite, hematite, and limonite are trans-opaque minerals, which means that as
209 the grain size decreases, the reflectance of these minerals decreases at wavelengths <550
210 nm and increases at wavelengths >550 nm (Hunt et al., 1971; Townsend, 1987). As the
211 grain size decreases, the absolute reflectance (brightness) of samples increases at
212 wavelengths >550 nm. On the other hand, the absorption depth at around 900 nm hardly
213 changed due to grain size variation, irrespective of the mineral forms of any iron oxide
214 present.

215 Spectral and chemical datasets were prepared in order to correlate the absorption depth
216 obtained from reflectance spectra in rocks with iron oxide abundance. A total of 43 rock
217 samples containing iron oxide minerals were used (hematite: 17, goethite: 21 jarosite: 5).
218 Of these, 14 rock samples containing goethite or hematite were selected from the data
219 catalog of Urai et al. (1989), and the rest of the samples were collected in Utah and
220 Nevada, USA. We used these data sets, because the range of Fe_2O_3 contents was much
221 larger than those of the seven powdered samples on Table 1.

222 The spectral dataset of Urai et al. (1989) consisted of reflectance spectra and chemical
223 and mineral compositions for 111 rock and mineral samples. Reflectance spectra were
224 measured in 800 contiguous spectral bands at wavelengths from 450 to 2600 nm using a
225 spectroradiometer (Infrared Intelligent Spectroradiometer, Geophysical Environment
226 Research, NY). The chemical compositions of the powdered samples were analyzed by
227 inductively coupled plasma - atomic emission spectroscopy (ICP-AES). Reflectance
228 spectra of the samples collected in Utah and Nevada, USA. were measured using a
229 spectroradiometer (FieldSpec pro, ASD Inc., CO) and their chemical compositions were
230 analyzed by XRF (ZSK Primus II instrument equipped with a Rh mode, Rigaku, Japan).
231 The depth of the absorption feature at around 900 nm due to ferric iron was calculated
232 from the reflectance spectra, and this was correlated with the actual Fe_2O_3 content (wt%)
233 in the chemical dataset.

234

235 2.4. Grain size regression

236 To account for the contribution of the grain size when estimating the iron oxide
237 abundance of trans-opaque minerals such as hematite and goethite, the difference between
238 the reflectance values at 550 nm and 760 nm (*Slope*) was used. This relationship can be

239 expressed as:

240

$$241 \quad \text{Slope} = 760R - 550R \quad (4)$$

242

243 where 760R and 550R are the reflectances at 760 and 550 nm, respectively; 760R is the
244 reflectance peak between the two absorption features centered at around 660 and 900 nm
245 due to ferric iron. In reflectance spectra of iron oxide minerals with various grain sizes,
246 550R behaves like a fixed point. Previous studies showed that the behavior of the spectral
247 changes of iron oxide minerals due to grain size differs at wavelengths shorter than 550
248 nm and longer than 550 nm (Hunt et al., 1971; Townsend, 1987). Thus, as the grain size
249 of iron oxide minerals increases, *Slope* increases as a function of the grain size (Fig. 2).
250 The effect of the grain size on a reflectance spectrum can therefore be suppressed by
251 subtracting *Slope* from the original spectrum:

252

$$253 \quad \text{Modified-R} = R - \text{Slope} \quad (5)$$

254

255 Here, Modified-R is a reflectance spectrum with grain size effect suppressed, and R is the
256 original reflectance spectrum. Modified-R can thus be used to suppress differences in
257 absolute reflectance due to the effect of grain size.

258

259 *2.5. Suppression of grain size and topography effects*

260 The conventional band ratio is an effective method for suppressing the effect of
261 topography, but the ratio of two reflectance values also depends on the original
262 reflectance values. For example, as grain size decreases, the reflectance increases and

263 consequently, the ratio of the two reflectance values decreases.

264

$$265 \quad \{R(\lambda_1) / R(\lambda_2)\}_{\text{large grain}} > \{R(\lambda_1) / R(\lambda_2)\}_{\text{small grain}} \quad (6)$$

266

267 where $R(\lambda_i)$ is the reflectance at wavelength λ_i .

268 Therefore, to suppress the effects of grain size and topography, it is necessary to
269 remove the effect of grain size prior to removal of the effect of topography by a
270 multiplicative operation (Fig. 3), as doing so will suppress changes in absolute reflectance
271 due to grain size.

272 To suppress the effects of grain size and topography, the continuum-removal method
273 must be applied to Modified-R. Thus, MCR-900D, which is the absorption depth at which
274 both effects are suppressed, can be obtained by the following equation:

275

$$276 \quad \text{MCR-900D} = 1.0 - \text{M900R} / \text{M900C} \quad (7)$$

277

278 Here, M900R and M900C are the spectral reflectance and its continuum at an absorption
279 center at which reflectance is the local minimum, respectively, derived from Modified-R.
280 An absorption center at around 900 nm is extracted in advance from the original
281 reflectance spectrum, because the continuum-removal method often changes the
282 wavelength of the absorption center (Haest et al., 2012). MCR-900D can be used to
283 effectively estimate iron oxide abundance (wt%) using spectral data, as effects of grain
284 size and topography suppressed.

285

286 *2.6. Application to AVIRIS*

287 *2.6.1. Study area*

288 The study area selected to evaluate the proposed method for estimating iron oxide
289 abundance was the Cuprite area in Nevada, USA, approximately 200 km northwest of
290 Las Vegas (the center of Fig. 5; N37.53, W117.18). The area has been selected by
291 geological remote sensing studies to evaluate new instruments and methods because of
292 the dry conditions, minimal vegetation, and abundance of clay and iron oxide minerals
293 due to extensive hydrothermal alteration (e.g., Clark et al., 2006; Swayze et al., 2014;
294 Yamaguchi and Naito, 2003). Hydrothermal alteration occurred as a result of interaction
295 between rocks and hot water, which was heated by magmatic activity in the mid- to late
296 Miocene (Ashley and Abrams, 1980). The area is divided into two subregions by
297 Highway 95, which bisects the study area along a north-south axis. In the western region,
298 Cambrian metasedimentary rocks, Paleogene conglomerate, and intrusive rocks that were
299 hydrothermally altered dominate, while in the east, Paleogene rhyolitic ash-flow tuff,
300 conglomerate, and basalt are distributed. Vegetation cover in this area is a few percent
301 and we assumed that vegetation did not markedly affect the spectral reflectance of the
302 AVIRIS data.

303

304 *2.6.2. Image data*

305 The proposed method for estimating the iron oxide abundance was applied to Airborne
306 Visible/Infrared Imaging Spectrometer (AVIRIS) image data. AVIRIS is a hyperspectral
307 sensor that has 224 contiguous spectral bands with wavelengths ranging from around 400
308 to 2500 nm and a spectral resolution ranging from 14 to 23 nm (Green et al., 1998).
309 AVIRIS is well suited to a wide variety of fields, including geology, environmental
310 studies, agriculture, and oceanography. The AVIRIS data for the Cuprite area was

311 obtained from the NASA website (http://aviris.jpl.nasa.gov/data/free_data.html). The
312 spatial resolution of the image is 17 m and the target area is about 8.7 km from east to
313 west and 10.4 km from north to south. We used the AVIRIS Standard Data Product of
314 Cuprite, which was converted to surface reflectance. We assumed that the atmospheric
315 correction was included in the reflectance conversion processing by NASA.

316

317 *2.6.3 Data processing*

318 The hydrothermally altered areas were also divided into silicified, opalized, and
319 argillized zones based on the degree of alteration (Ashley and Abrams, 1980). The
320 silicified zone consists mainly of quartz and chalcedony, the opalized zone has alunite
321 and kaolinite deposits, and the argillized zone has kaolinite deposits (Swayze et al., 2014).
322 A supervised classification of AVIRIS VNIR data (400-1350 nm) was conducted in order
323 to comprehend the distribution of various iron oxide minerals and to select an appropriate
324 linear regression for each pixel.

325 Iron oxide abundance in the Cuprite area was estimated using AVIRIS data after
326 calculating the absorption depth by MCR-900D. Absolute values of the iron oxide
327 abundance for each pixel were determined using the three linear regressions
328 corresponding to the different mineral forms of iron oxide.

329

330 **3. Results**

331 *3.1. Verification of spectral measurement results*

332 The average difference between the maximum and minimum values of 900D in the
333 reflectance spectra of seven different grain size samples was less than 0.01% (reflectance
334 value). We can therefore conclude that the absorption depth is almost constant even if the

335 grain size changes. In previous studies, the absorption depth at around 900 nm of goethite
336 and hematite increased as the grain size decreased (Hunt et al, 1971), while the absorption
337 depth at around 1000 nm of the other minerals such as amphibole and olivine became
338 shallower as the grain size decreased (Hunt and Salisbury, 1970); however, such a trend
339 was not apparent in our measurement results. In addition, the differences among the
340 reflectance spectra of different grain size samples were almost constant (less than 1%
341 reflectance difference), which means that neither the shape of the spectra nor 900D
342 changed, and that grain size has an additive effect on reflectance spectra. There were
343 differences between our measurements and the measurements by Hunt and Salisbury
344 (1970) and Hunt et al. (1971). For example, we used rock powder samples containing iron
345 oxides, while Hunt and Salisbury (1970) and Hunt et al. (1971) used powders of pure
346 minerals. Moreover, we measured samples measuring 25 to 500 μm , while the
347 measurements of Hunt and Salisbury (1970) and Hunt et al. (1971) used very fine grain
348 samples (0-5 and 0-74 μm). In addition, the absorption depths due to OH molecular
349 vibrations at around 1420 nm, 1920 nm and 2200 nm did not change markedly for
350 different grain sizes.

351 The reason why reflectance increases while the spectral shape and absorption depth
352 remain relatively constant as the grain size decreases is probably due to the increase in
353 the amount of first surface reflection and multiple scattering (Cooper and Mustard, 1999).
354 Surface scattering is dominant in the VNIR regions, while volume scattering is dominant
355 in the thermal infrared region. The reflectance from a single grain is generally determined
356 by boundary reflection (first surface reflection), internal reflection, effective grain size,
357 and the absorption coefficient (Hiroi and Pieters, 1992). These reflections are not affected
358 by wavelength because the grain sizes examined in this study are larger than the

359 wavelengths of VNIR light. In addition, since surface backscattering increases as grain
360 size decreases, reflectance increases independently of wavelength. On the other hand,
361 previous studies demonstrated an increase in absorption depth with decreasing grain size
362 until the absorption path length coincided with the mean optical path length (Hapke,
363 1993). The absorption depth at around 600 nm decreased with decreasing grain size
364 because volume scattering played a smaller role compared to first surface scattering.
365 However, at around 900 nm, the absorption depths were almost constant in this study. It
366 is considered that the absorption coefficient did not change markedly with grain size,
367 because larger grains might consist of aggregations of finer grains in the actual rock and
368 sand samples. It should be noted that grain shape also plays an important role in the
369 effective grain size and in changes in the scattering intensity (Hiroi and Pieters, 1992).

370

371 *3.2. Correlation between absorption depth and iron oxide abundance*

372 The depth due to ferric iron at the absorption center derived using MCR-900D was
373 correlated with the actual Fe₂O₃ content in the spectral and chemical datasets of the 43
374 rock samples to estimate surface iron oxide abundance in remotely sensed image data.
375 MCR-900D was positively correlated with Fe₂O₃ content ($R^2 = 0.49$) with a root mean
376 squared error (RMSE) of 3.57 wt% using a linear regression for all the iron oxide samples.
377 The variation in the mineral forms of iron oxide caused differences in the brightness,
378 shape, and position of absorption centers in their reflectance spectra in the VNIR region.
379 For example, the absorption center in the reflectance spectra of hematite, goethite and
380 jarosite is around 850 nm, 940 nm and 900 nm (Hunt and Ashley, 1979).

381 Figure 4 shows the correlations between MCR-900D and Fe₂O₃ content for the
382 different mineral forms of iron oxide with significantly higher correlation coefficients

383 (hematite: $R^2 = 0.75$, goethite: $R^2 = 0.82$, jarosite: $R^2 = 0.94$). Moreover, RMSEs obtained
384 using linear regression were highly improved for each mineral form (hematite: 0.82 wt%,
385 goethite: 2.59 wt%, jarosite: 0.31 wt%). The major reason for the deviations from the
386 regression lines was likely due to the different compositions of the various minerals in
387 the samples which may have caused different multiple reflections. After classifying the
388 mineral forms of iron oxide using the wavelengths of the absorption center, iron oxide
389 abundance was estimated by image processing using the following three linear
390 regressions (hematite: $y = 5.63x - 0.02$, goethite: $y = 25.82x - 1.13$, jarosite: $y = 20.84x$
391 $- 2.27$; where MCR-900D is substituted for x and y is iron oxide abundance). In this study,
392 each pixel was classified according to the wavelength of absorption center; <900 nm to
393 hematite, 900 - 930 nm to jarosite, and >930 nm to goethite (Fig. 3).

394

395 *3.3. AVIRIS data analysis*

396 The supervised classification result (Fig. 5) showed that iron oxide minerals were
397 correlated with the opalized and argillized zones. Moreover, the mineral forms of iron
398 oxide differed between the western and eastern regions. Jarosite and jarosite + goethite
399 are common in the western region, while hematite is locally distributed in the eastern
400 region and to the north of the silicified zone. Visually, the color of the surface rocks
401 including iron oxide minerals is yellowish in the western region and red in the eastern
402 region. Differences in the distribution of iron oxides probably depend on the depth of
403 erosion and it is suggested that the western region is more eroded than the eastern region.

404 In the iron oxide abundance image (Fig. 6), regions with larger absorption depths at
405 around 900 nm in the hydrothermally altered areas were well represented by all three
406 methods. The highest level of iron oxide abundance was approximately 6 wt%, and such

407 areas were distributed mainly in the western hills of the Cuprite area (shown as red in Fig.
408 6). As can be seen from iron oxide abundance map and the mineral distribution map (Fig.
409 5), in terms of iron oxide abundance, the distribution of hematite (ca. 2 wt%) in the eastern
410 area was relatively low compared to that in the western area, because the slope of the
411 regression line in hematite samples was smaller than those of jarosite and goethite (Fig.
412 4). It is suggested that rocks with hematite appear reddish, even if iron oxide abundance
413 is low. In addition, though iron oxide was not extracted from the silicified zone, the
414 enrichment zones were consistent with the opalized and argillized zones. Enrichment
415 zones consisting of iron oxide with jarosite and goethite are important in exploration
416 geology settings.

417

418 *3.4. Validation of AVIRIS results*

419 Although the results obtained using AVIRIS data processed by these three methods
420 (MCR-900D, CR-900D, and 900D) were generally similar (Fig. 7), the results obtained
421 by MCR-900D showed a strong contrast between the altered and non-altered areas.
422 Southwest of the hills in the western region, absorption depth estimated by CR-900D was
423 larger than that estimated by the other methods because the reflectance was lower than
424 the altered areas in the hills. Conversely, reflectance was lower in the areas with high
425 levels of hematite in the eastern hills. In addition, iron oxide abundance estimated by
426 900D and CR-900D in the hills to the west and southwest of the western region was higher
427 than that estimated by MCR-900D. It is therefore necessary to analyze the chemical
428 compositions in the surface rocks in these regions in order to verify the abundance maps.

429 To investigate whether the effect of topography was suppressed by the processing
430 methods, the average reflectance for all spectral bands was correlated with 900D and

431 MCR-900D, respectively. It was considered that 900D would essentially only reflect the
432 iron oxide abundance, and that 900D would be independent of the average reflectance. In
433 other words, if the correlation between the average reflectance was weak, then we could
434 assume that the effect of topography was suppressed. However, our results showed that
435 the average reflectance for all bands decreased as the topography effect increased (i.e., an
436 increase in the topographic slope), and the values obtained by 900D decreased
437 accordingly. By being affected by topography, 900D is considered to be dependent on the
438 average reflectance for all bands, implying that iron oxide abundance would be correlated
439 with the average reflectance for all bands.

440 The coefficients of determination (R^2) between the average reflectance obtained for all
441 AVIRIS bands and 900D and MCR-900D for the Cuprite area were 0.47 and 0.28,
442 respectively; 900D and average reflectance showed a higher correlation compared to
443 MCR-900D, indicating that 900D did not suppress the effect of topography. However,
444 MCR-900D and average reflectance were less strongly correlated, which means that
445 MCR-900D is not affected by topography, likely due to the continuum-removal method.
446 MCR-900D is thus effective for estimating iron oxide abundance by using reflectance
447 spectra in remotely sensed data, as it suppresses the effects of grain size and topography.

448

449 **4. Discussion**

450 Generally, the SiO_2 content of rocks is inversely proportional to the Fe_2O_3 content.
451 SiO_2 -rich rock samples contain higher amounts of transparent minerals such as quartz,
452 and the high reflectance associated with SiO_2 -rich sediments is strongly affected by these
453 minerals. Consequently, the increase in the optical depth attributed to these transparent
454 minerals results in overestimates of Fe_2O_3 abundance in Fe_2O_3 -poor and SiO_2 -rich

455 samples (Haest et al., 2012). Though SiO₂-rich samples contain lower amounts of Fe₂O₃,
456 no such trend was observed in this study. In addition, aluminum substitution in hematite
457 and goethite brings about a shift in the absorption center due to ferric iron at around 900
458 nm toward a longer wavelength (Buckingham and Sommer, 1983; Sommer and
459 Buckingham, 1981). However, since the Al₂O₃ content in rock samples did not affect the
460 absorption depth of iron oxide in this study, it is considered that the presence of other
461 chemicals and their abundances did not alter the absorption depth of ferric iron.

462 Another factor that might affect the correlation between MCR-900D and Fe₂O₃
463 content is the difference between the spectral measurements and chemical analysis of
464 each sample. In the dataset of Urai et al. (1989), the reflectance spectra were measured
465 using the natural surface of samples without any processing. Consequently, mineral and
466 chemical compositions are restricted to the surface of samples. On the other hand, for the
467 chemical analysis in this study, each sample was crushed and powdered, which means the
468 results of the chemical analysis reflected the bulk chemical composition of a sample, i.e.,
469 not only the surface of the sample but also its internal composition. It is therefore likely
470 that the variation observed between the absorption depth and the Fe₂O₃ content occurred
471 due to the difference of the measured materials.

472 The 900D data obtained for the different mineral forms of iron oxide were positively
473 correlated with Fe₂O₃ content (hematite: R² = 0.46, goethite: R² = 0.39, jarosite: R² =
474 0.87). In addition, the RMSE obtained from the linear regression was 1.20 wt% in the
475 hematite samples, 4.70 wt% in goethite samples, and 0.47 wt% in jarosite samples. Thus,
476 MCR-900D was more strongly correlated with iron oxide abundance. The improvement
477 in the correlation observed between MCR-900D and Fe₂O₃ content, compared with 900D
478 and Fe₂O₃ content, was due to enhancement or exaggeration of the absorption feature at

479 around 900 nm by the continuum-removal method. In addition, these spectral data were
480 all measured in the laboratory, and it is considered that in the field, satellite or aircraft
481 imagery processed with 900D would be more affected by topography compared to MCR-
482 900D, because 900D would be affected by the topography effect as well as noise in actual
483 sensor measurements.

484 Significantly high correlations were observed between CR-900D and the different
485 mineral forms of Fe_2O_3 in the samples (hematite: $R^2 = 0.71$, goethite: $R^2 = 0.75$, jarosite:
486 $R^2 = 0.94$). Also, the RMSE obtained by linear regression was 0.88 wt% in the hematite
487 samples, 3.07 wt% in the goethite samples, and 0.32 wt% in the jarosite samples.
488 Compared to CR-900D, MCR-900D showed improved correlations in the estimation of
489 iron oxide abundance. As CR-900D is not affected by the effect of topography, we
490 attributed the improvement observed using MCR-900D to suppression of the effect of
491 grain size (the variation of absolute reflectance), obtained by subtracting *Slope* from the
492 spectral reflectance before continuum-removal processing. In jarosite samples, the
493 absence of grain size variation meant that the correlations obtained with MCR-900D did
494 not show any improvement compared to CR-900D, because the grain size of all jarosite
495 samples was 180-500 nm. The correlation between CR-900D and average reflectance in
496 the visible wavelength region ($R^2 = 0.52$) was stronger than that between MCR-900D and
497 average reflectance ($R^2 = 0.39$). This means that CR-900D is still affected by the absolute
498 reflectance at the absorption center, which is affected by chemical and mineral
499 composition as well as grain size variation, because the continuum-removed spectrum is
500 derived from the ratio between the original spectrum and its continuum. Consequently,
501 CR-900D is lower for rocks compared to material with a smaller grain size and higher
502 reflectance.

503 Compared to CR-900D, the difference in estimates of iron oxide abundance obtained
504 by MCR-900D, which is an estimate of the difference between the estimated maximum
505 and minimum abundance values in reflectance spectra obtained for different grain size
506 samples was smaller. However, some samples did not show an improvement in
507 suppression of grain size effect by MCR-900D. The potential effectiveness of MCR-900D
508 is based upon the assumptions that *Slope* increases linearly as grain size decreases, and
509 that reflectance at 550 nm is independent of variations in grain size. If these conditions
510 are not met, then MCR-900D cannot suppress the effect of grain size effectively. On the
511 other hand, the difference in estimates of iron oxide abundance by CR-900D might be
512 larger if the range of grain sizes increases beyond the range examined in this study (<25-
513 500 μm).

514 In comparisons among 900D, CR-900D and MCR-900D, the most accurate estimation
515 of iron oxide abundance on rock surfaces was obtained by MCR-900D, which suppressed
516 the effects of grain size and topography. However, it should be noted that MCR-900D
517 (*Slope*) is dependent upon the spectral properties of trans-opaque minerals, that is, iron
518 oxide minerals. MCR-900D is thus not suitable for estimating the abundance of other
519 chemicals/minerals on rock surfaces with transparent or opaque minerals, such as those
520 with clay minerals at around 2200 nm.

521

522 **5. Summary and Conclusions**

523 The authors developed a new method, called MCR-900D, for estimating iron oxide
524 abundance in rocks containing iron oxide minerals. The method is capable of suppressing
525 the effects of grain size and topography in estimates of iron oxide abundance, which is
526 related to the depth of absorption features conferred by the presence of ferric iron at

527 around 900 nm (900D).

528 First, to clarify the effect of grain size on reflectance spectra of rock samples containing
529 iron oxide minerals, spectra were measured in samples of different grain sizes (<25 μm ,
530 25-45 μm , 45-75 μm , 75-125 μm , 125-180 μm , 180-250 μm , and 250-500 μm) using a
531 spectroradiometer. The effects of grain size were:

532 (1) The spectral reflectance increased as the grain size decreased regardless of the
533 chemical/mineral compositions at wavelengths longer than 550 nm.

534 (2) The shape and the absorption depth for ferric iron absorption spectra were almost
535 constant.

536 (3) The difference between reflectance at 550 nm and 760 nm in rocks containing iron
537 oxide minerals (*Slope*) was a function of grain size.

538 These characteristics were considered in developing MCR-900D in order to suppress the
539 effect of grain size.

540 On the other hand, since the spectral reflectance of minerals and rocks is
541 multiplicatively affected by topographic slope, the continuum-removal method was
542 employed in pre-processing to suppress the topography effect as well as enhance the
543 absorption features. In order to suppress the effects of grain size and topography, the
544 effect of grain size needed to be removed by a subtractive operation prior to removal of
545 the effect of topography using a multiplicative operation. By performing *Slope* and then
546 continuum-removal processing in turn, MCR-900D can be used to estimate the absorption
547 depth while suppressing the effects of grain size and topography.

548 MCR-900D was then applied to samples in spectral and chemical datasets with known
549 Fe_2O_3 contents. Analysis of 43 rock samples revealed the following:

550 (1) The mineral forms of samples need to be considered when extracting spectral

551 information about the wavelength of the absorption center as well as the absorption
552 depth. The correlations between MCR-900D and the Fe_2O_3 content were
553 significantly improved (all samples: $R^2 = 0.49$, hematite: $R^2 = 0.75$, goethite: $R^2 =$
554 0.82 , jarosite: $R^2 = 0.94$).

555 (2) The presence of other chemicals, such as SiO_2 , did not affect the correlation between
556 MCR-900D and the actual Fe_2O_3 content.

557 (3) The results of processing using only the continuum-removal method depended on
558 the original absolute reflectance, which meant that the grain size effect cannot be
559 suppressed when using the continuum-removal method alone.

560 Finally, MCR-900D was applied to the AVIRIS data for the Cuprite area in Nevada,
561 USA. The absorption depth and iron oxide abundance map using linear regression well
562 represented the geology, particularly the distributions of the hydrothermally altered areas.
563 Assuming from the correlations with the average reflectance for all bands, it was
564 concluded that MCR-900D successfully suppressed the topography effect. MCR-900D is
565 thus a simple and effective method for estimating the iron oxide abundance in rocks that
566 can provide valuable information on hydrothermal alteration as well as exploration of ore
567 deposits.

568

569

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579

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671 Captions of Table and Figures

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679

680 Figure 2. *Slope* in reflectance spectra of different grain size samples. *Slope* decreases as
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682

683 Figure 3. Processing sequence of the proposed method for different forms of iron oxide
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685

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687 (b) goethite (n=21) and (c) jarosite (n=5) samples.

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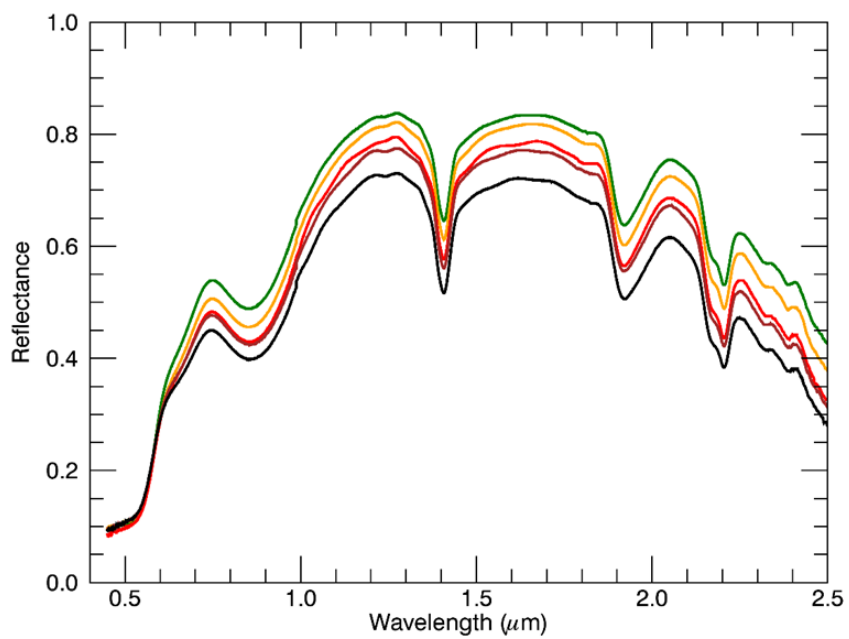
699

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 701 F, G) and sand (C, D) samples used to investigate the grain-size effect.

Sample	Form	SiO ₂ wt%	Al ₂ O ₃ wt%	Fe ₂ O ₃ wt%	CaO wt%	MgO wt%	Na ₂ O wt%	K ₂ O wt%	LOI wt%	Total wt%
A	hematite	73.20	12.30	3.73	0.55	0.50	1.08	5.02	4.50	100.88
B	hematite	75.90	10.95	3.84	1.36	0.19	0.08	0.15	7.68	100.15
C	hematite	82.69	5.61	1.01	0.69	0.98	0.79	2.09	2.86	96.72
D	hematite	96.63	0.57	0.06	0.00	0.04	0.00	0.32	0.40	98.01
E	goethite	85.90	1.98	8.10	0.10	0.07	0.04	0.16	2.28	98.63
F	goethite	74.00	12.85	4.03	0.14	0.26	0.38	1.91	8.05	101.62
G	jarosite	37.10	22.00	3.59	0.37	0.05	1.05	4.52	28.10	96.78

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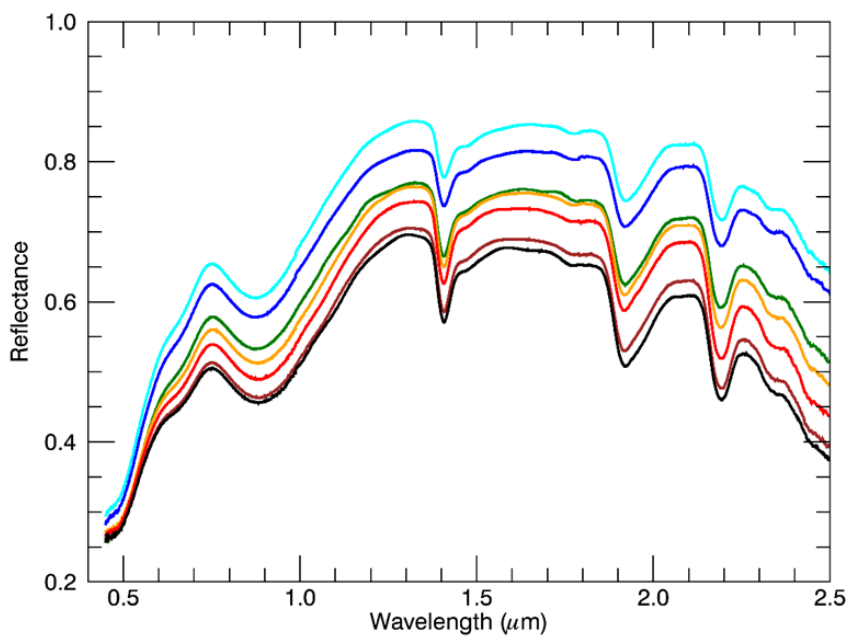
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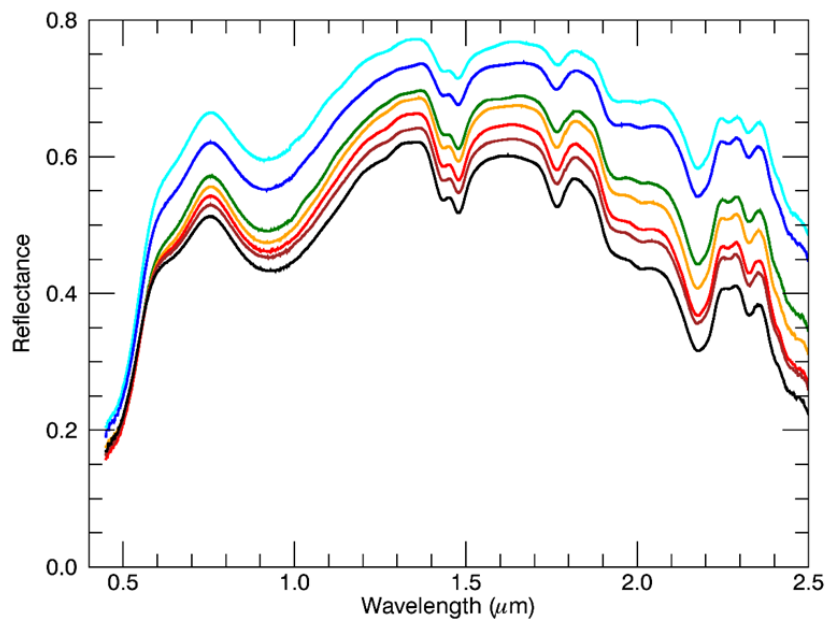
(a) Sample B



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(b) Sample F



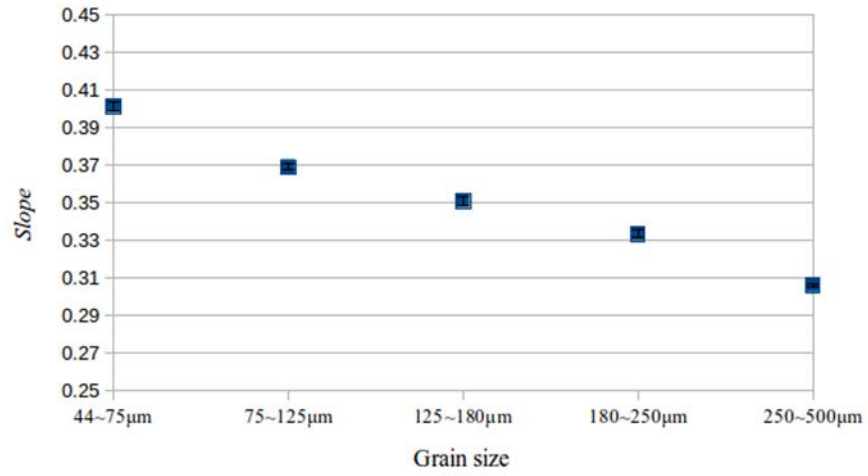
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(c) Sample G

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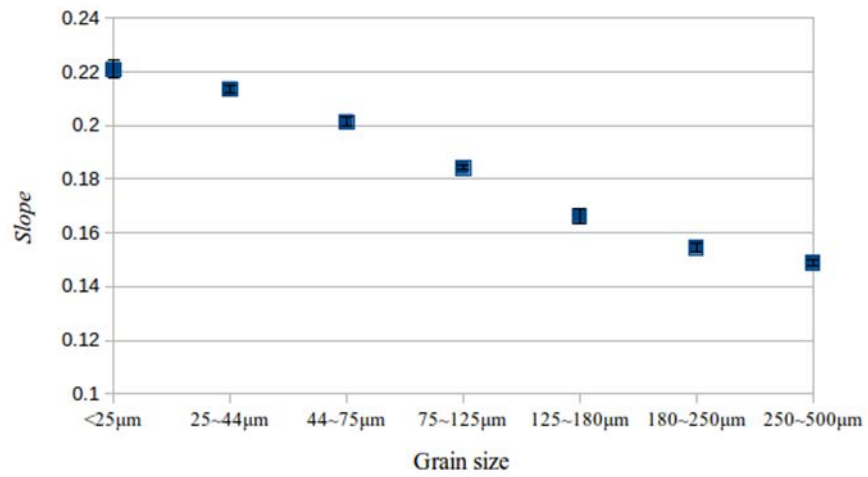


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(a) Sample B

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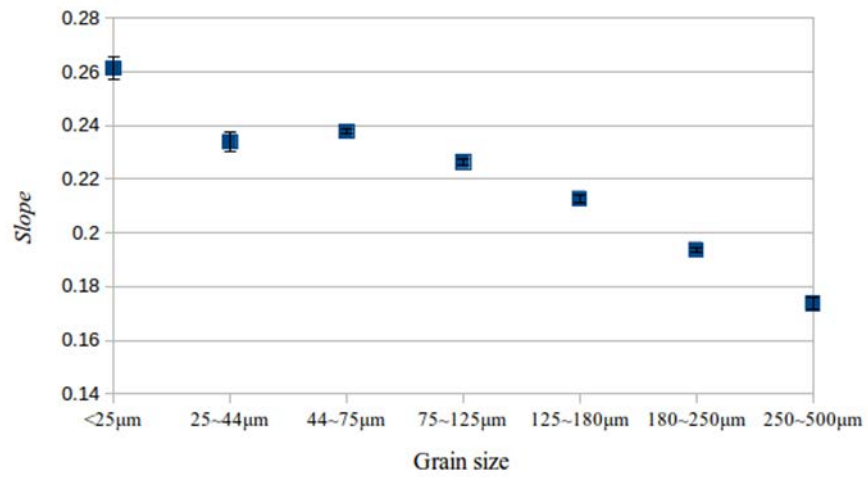


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(b) Sample F

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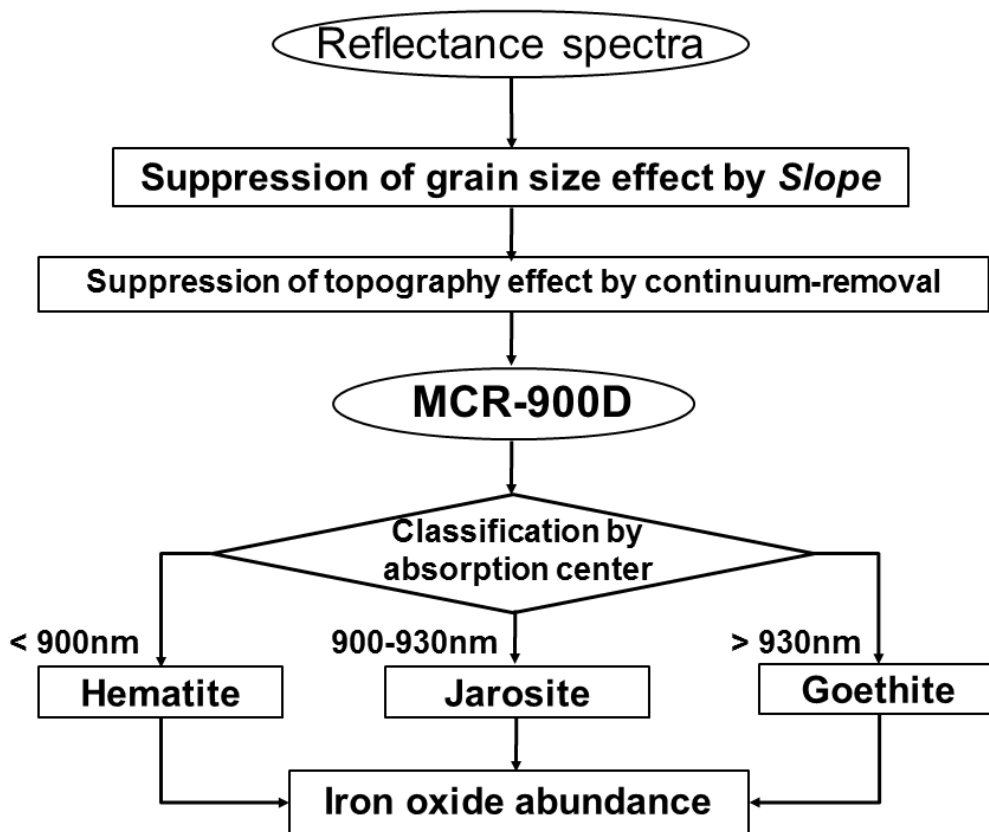
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(c) Sample G

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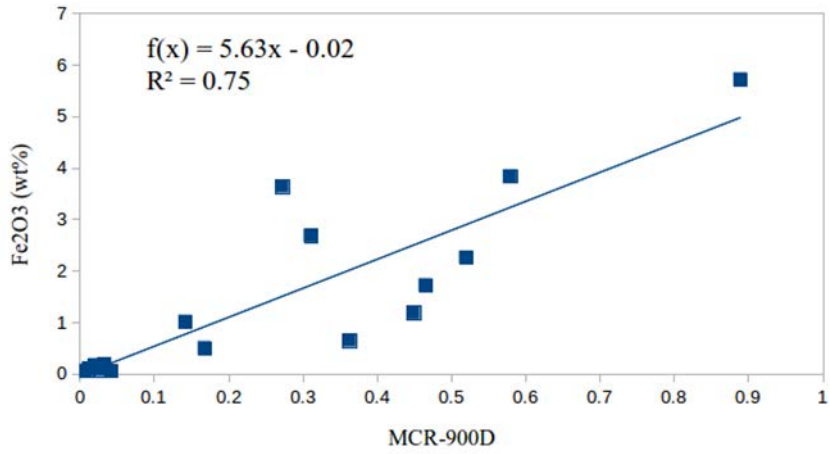
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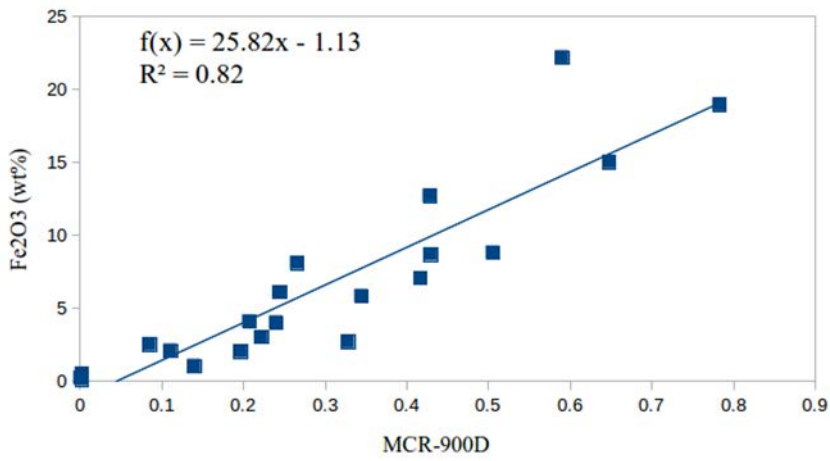


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(a) Hematite samples

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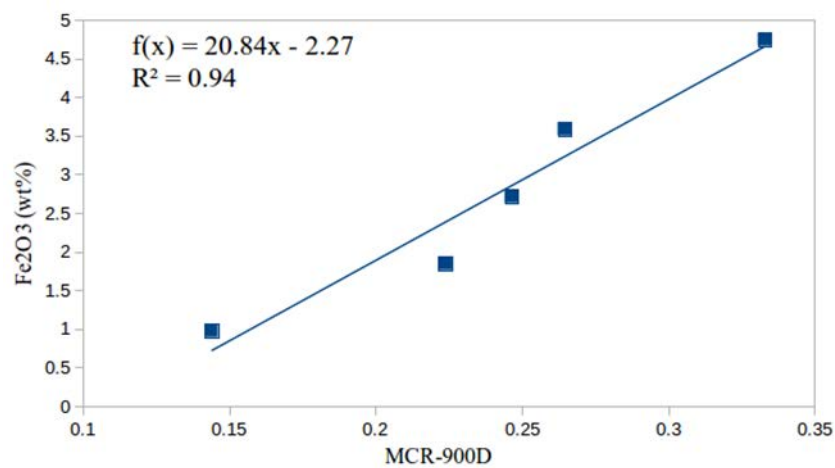


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(b) Goethite samples

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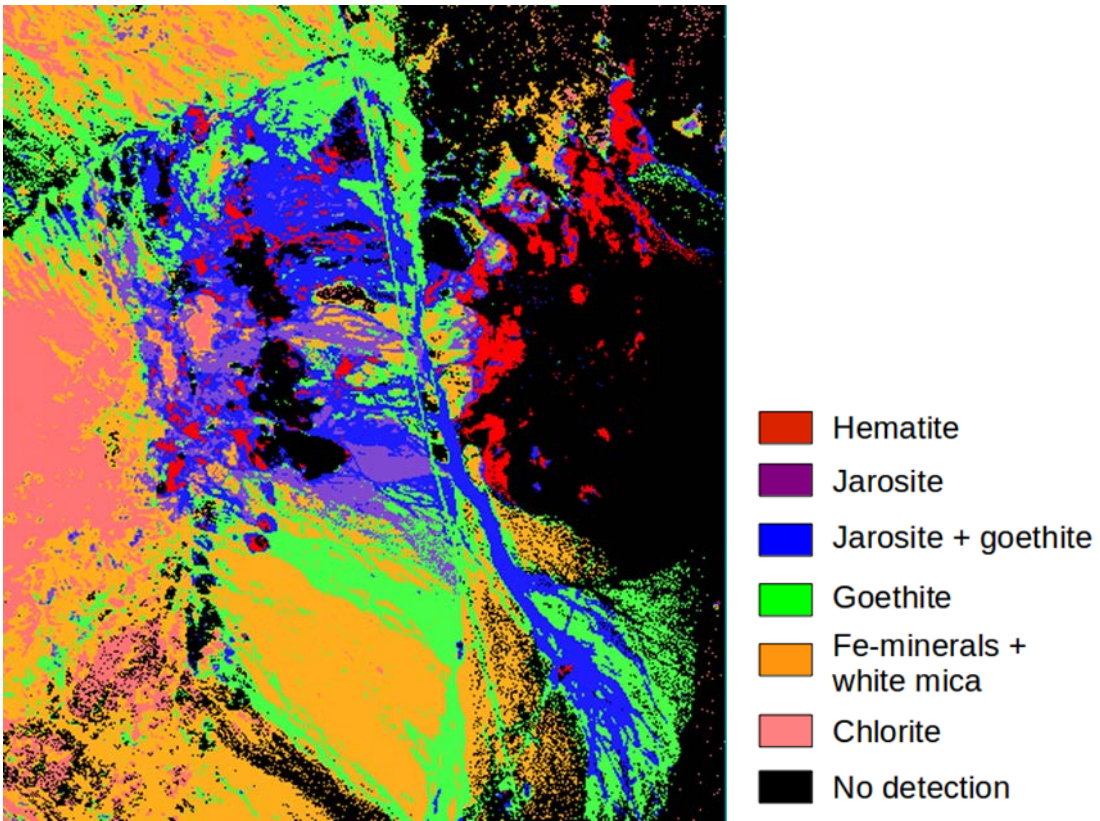
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(c) Jarosite samples

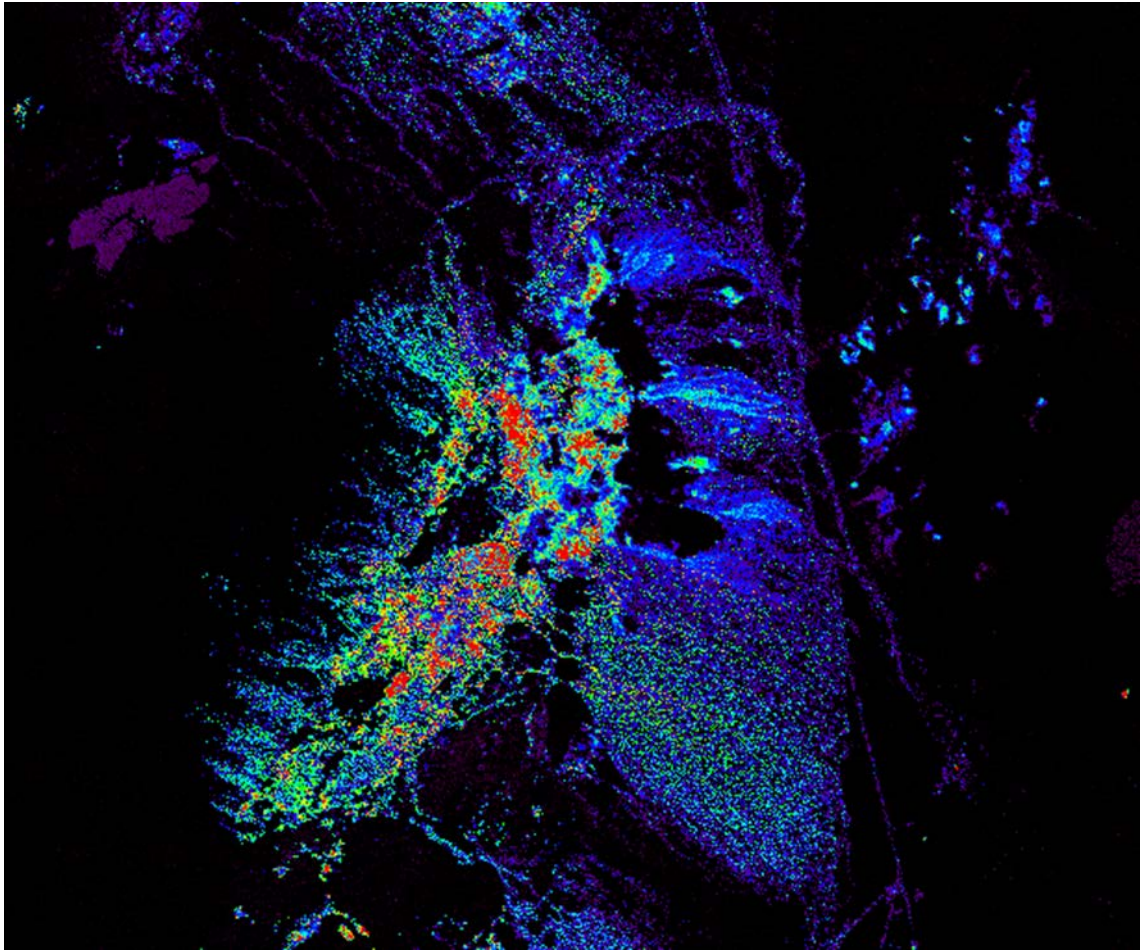
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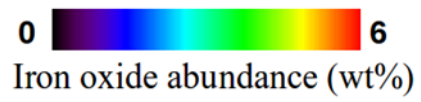
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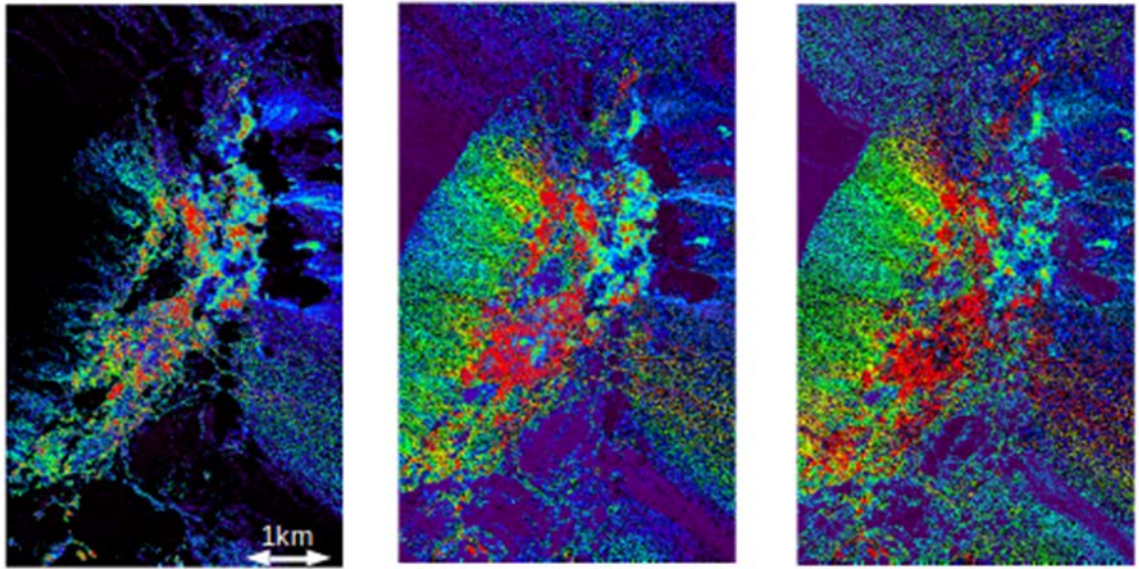


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