

1 **Vertical wavenumber spectral characteristics of temperature in the**
2 **stratosphere-mesosphere over tropical and subtropical regions**

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7
8 **Abstract**

9 The vertical wavenumber spectra over tropical location, Gadanki (13.5° N, 79.2° E) and sub-
10 tropical location, Mt. Abu (24.5° N, 72.7° E) is studied using the temperature measurements from
11 ground based Rayleigh Lidar and space borne satellite observations. The slope values are lesser
12 over Gadanki than at Mt. Abu for almost all the altitudes except for 40-50 km where it is nearly
13 same and 60-70 km exhibiting opposite nature. Unusual spectral slope of -6.97 (Mt. Abu) and -
14 0.09 (Gadanki) is seen at the altitude of 40-50 km in satellite temperature. Characteristics of wave
15 oscillations perceived over both the stations are described.

16 Key words: lidar; vertical wavenumber spectra; spectral slope; wave oscillation

17
18 **1. Introduction**

19 Atmospheric gravity waves (GWs) are abundantly generated in the earth's lower atmosphere
20 which affects the energy and momentum budget in the middle and upper atmosphere by regulating
21 and altering the thermodynamical structure of the atmosphere significantly (Fritts and Alexander,
22 2003). Therefore, it is essential to have a thorough knowledge of the atmosphere GWs
23 characteristics for better understanding of the global mesospheric dynamics. When the GWs
24 proliferate vertically upward, their amplitude intensifies exponentially due to decrease in
25 atmospheric density resulting in convective or dynamic instability in the surrounding area.
26 Therefore, the GWs releases energy to maintain overall stability, during upward propagation, while
27 the power of wavenumber spectra and the spectral slope becomes saturated. VanZandt (1982)
28 coined the purported 'universal' wavenumber spectra referring to the independent behavior of the

29 frequency, horizontal and vertical wavenumber spectra of horizontal wind and temperature
30 fluctuations (specifically in the higher wavenumber region) irrespective of altitude, season and
31 geographical location. Numerous studies have been show similarities or dissimilarities in
32 observation with VanZandt (1982) proposition using ground-based and satellite-borne instruments
33 (e.g., Eckermann, 1995; Ghosh et al., 2018; Larsen et al., 1986; Smith et al., 1987; Zhang et al.,
34 2017a,b and references therein). Gardner et al., (1993) derived a model of vertical wavenumber
35 (VWN) spectra of GWs, for atmospheric density and velocity fluctuations, where the spectral
36 slopes are ‘2’, ‘-3’ and ‘-5/3’ for unsaturated, saturated and turbulence region respectively.
37 Recently, Ghosh et al., (2018) found the saturated spectral slope of vertical wind, observed using
38 mesosphere stratosphere troposphere (MST) radar over a tropical location Gadanki (13.5° N, 79.2°
39 E), to be ~ -6 , ~ -9 and ~ -12 in some cases. The foremost purpose of this paper is to investigate the
40 VWN spectral characteristics over a tropical location Gadanki (13.5° N, 79.2° E) and a sub-tropical
41 location Mt. Abu (24.5° N, 72.7° E). For this work, simultaneous observations of ground-based
42 Rayleigh lidar at both the locations are used along with Sounding of the Atmosphere using
43 Broadband Emission Radiometry (SABER) temperature measurements onboard Thermosphere
44 Ionosphere Mesosphere Energetics and Dynamics (TIMED) satellite.

45 **2. Data and instrumentation**

46 Rayleigh lidar at Mt. Abu (24.5° N, 72.7° E) was installed by Physical Research Laboratory (PRL)
47 and at Gadanki (13.5° N, 79.2° E) by National Atmospheric Research Laboratory (NARL) for
48 middle atmospheric temperature measurements. The details of both the lidar systems are available
49 in Sharma et al., (2017) and Sivakumar et al., (2003). For the present study, the lidar temperature
50 covers the altitude of 33-70 km divided into 33-40 km, 40-50 km, 50-60 km and 60-70 km over
51 Mt. Abu (during 14-21 March 2004) and 30-66 km divided into 30-40 km, 40-50 km, 50-60 km
52 and 60-66 km over Gadanki (during 19-26 March 2004). SABER is one of National Aeronautics
53 and Space Administration’s TIMED satellite with latitude coverage of 52° N to 83° S (all through
54 the north observing mode for nearly 60 days) or 52° S to 83° N (in the south observing mode)
55 (Russell III et al., 1999). Therefore, the TIMED satellite can continuously access the latitudes 52°
56 S to 52° N (Mertens, 2004; Mertens et al., 2001). The temperature measurements from SABER
57 instrument is obtained for the longitude grid of 23-25° N (12-14° N) and longitude grid of 70-73°
58 E (78-81° E) over Mt. Abu (Gadanki) during the time period of 20:00 UT to 21:30 UTC

59 respectively. Earlier studies reported that the spectra are ‘red’ at the higher wavenumbers for most
60 of the atmospheric variables (Blackman and Turkey, 2009; Percival and Walden, 1993). Therefore,
61 various methods are used in the past to overcome or reduce spectral leakage. To conserve the
62 ‘spectral density’ at small wavenumbers, it is found that the ‘prewhitening with first differences’
63 methodology is very efficient (Nastrom and VanZandt, 2001; Tsuda et al., 1989). The daily
64 temperature profiles are normalized (T'/\bar{T} , where T' is the temperature fluctuation and \bar{T} is the
65 mean temperature) (Allen and Vincent, 1995; Guharay and Sekar, 2011; Nastrom et al., 1997;
66 Tsuda and Hocke, 2002; Yan et al., 2019) and subjected to spectral analysis using Fast Fourier
67 Transform (FFT) (after ‘prewhitening/postcoloring’ techniques) to acquire the proliferation
68 characteristics of dominant wave oscillations and the VWN spectral slope values. The normalized
69 daily temperature data is ‘prewhitened’ by removing the ‘first difference’ of the residuals (X_q)
70 using the methodology suggested by Nastrom and VanZandt, (2001) as follows:

$$71 \quad \bar{X}_q = X_q - \alpha X_{q-1} \quad (1)$$

72 Here, the value of α is 1 (Dewan and Grossbard, 2000; Nastrom and VanZandt, 2001; and
73 references therein for more details of 'prewhitening').

74 Using Fast Fourier Transform (FFT) analysis, the spectral amplitude of the ‘prewhitened’ data is
75 determined. The ‘postcolored’ spectra (F_{PWPC}) is obtained using the following equation (Nastrom
76 and VanZandt, 2001):

$$77 \quad F_{PWPC}(m_{r,j}) = \frac{F_{PW}(m_{r,j})}{2[1-\cos(2\pi j/M)]} \quad (2)$$

78 where $m_{r,j} \rightarrow j$ th wavenumber in radians/meter ($m = m_r/2\pi$) in cycles/meter, and $M \rightarrow$ total
79 number of data points.

80 In order to show the difference in spectral slope before ‘prewhitening/postcoloring’, the power
81 spectral density (PSD) obtained using Fast Fourier Transform (FFT) on the normalized
82 temperature data without ‘prewhitening/postcoloring’ is also calculated and both the slope values
83 (where ‘Slope 1’ indicates slope values without ‘prewhitening/postcoloring’ and ‘Slope 2’ denotes
84 slope values after inclusion of ‘prewhitening/postcoloring’) are tabulated in **Table 1** and also
85 plotted in all the figures (**Figure 1-9**). To measure the accuracy of VWN spectral slopes, co-
86 efficient of determination (R^2) of the spectral slopes using lidar and SABER/TIMED observation
87 [with and without Pre-whitening and Post-Coloring (PWPC)] over Mount Abu (24.5° N, 72.7° E)
88 and Gadanki (13.5° N, 79.2° E) for the individual height regions covering the whole range of 30-

89 70 km is calculated (**Table 2 and 3**; where **Table 2** is based on lidar and **Table 3** depicts the
90 SABER/TIMED observation on the nearest possible day over both stations depending on data
91 availability). The standard deviation of R^2 values is indicated with ‘ \pm ’ sign.

92 **3. Results and discussion**

93 For the PSD calculation, nightly mean data has been used and the selected nights have more than
94 2-3 hours observation on each day (depending upon local weather conditions). Each profile is of
95 5 or 10 minutes and all the profiles collected in a single night are integrated to improve the signal
96 to noise ratio. Altitude profile of VWN spectra using lidar temperature observation over Mt. Abu
97 during 14-17 March 2004 shown in **Figure 1** where the x-axis denotes the natural logarithm of
98 wavenumber (cycles/km) and y-axis denotes the natural logarithm of PSD (K^2 /wavenumber). It
99 is to be noted that all the slope values discussed here are approximate. It is observed that the
100 spectral slope is more negative in the altitude range of 40-50 km (fluctuating between -2.67 to -
101 2.82) and 60-70 km (fluctuating between -2.58 to -2.72) except for 16 March where the slope is
102 most negative (-2.88) at the height of 56-60 km. In case of the altitude profile of VWN spectra
103 during 18-21 March 2004 (**Figure 2**), similar pattern is observed with most negative slopes in the
104 elevations of 40-50 km (between -2.5 to -2.94) and 60-70 km (between -2.68 to -2.79) except for
105 18 March which shows the spectral slope of -2.79 at the height of 50-60 km. Ghosh et al., (2018)
106 suggested that the tropopause may act as the source of new gravity waves resulting in positive
107 VWN spectral slope value. Since the stratopause exists in the altitude of 45-50 km, the more
108 negative slope at that elevation can be contributed to the presence of stratopause which is a
109 transition zone of atmosphere. It is to be also taken into account that the station Mt. Abu is located
110 at an elevation of about 1.7 km from the mean sea level which may also add up to the changes in
111 spectral slope due to the presence of terrain generated gravity waves.

112 In case of the altitude profile of VWN spectra over Gadanki during 19-22 March 2004 (**Figure**
113 **3**), it is perceived that the spectral slope is more negative at the heights of 30-40 km (ranging
114 between -2.71 to -3.01) and 40-50 km (between -2.94 to -3.25) in contrast to the observations over
115 Mt. Abu. Although the spectral slope values in the elevations of 50-60 km are more negative than
116 those of 60-66 km, still they less negative than the ones in 30-40 km and 40-50 km heights. Similar
117 pattern is detected for the VWN spectral slope values in **Figure 4** except for 26 March 2004 where
118 the most negative slopes occur at 40-50 km and 50-60 km. Guharay and Sekar, (2011) used

119 Rayleigh lidar observations over Gadanki operated in cloud free nights on a systematic manner
120 with more than 4 hours of unswerving observation to investigate GWs of periodicities less than 4
121 hours (during March 1998 – December 2008). They reported that the spectral slope in the upper
122 stratosphere (35-50 km) is considerably higher (~ -2.83) and nearly corresponds to the theoretical
123 magnitude of spectral slope obtained than those observed in the mesospheric (50-72 km) heights
124 (~ -2.53). They also suggested that this may affect the mesosphere such that it becomes more
125 susceptible to maintain ‘nonlinear wave interaction processes’ rather than ‘individual wave packet
126 propagation’. In the present case, the spectral slope values over Mt. Abu are below ‘-3’ ranging
127 from -2.14 to -2.93 (without ‘prewhitening/postcoloring’) and -2.71 to -2.97 (with
128 ‘prewhitening/postcoloring’) (during 14-17 March 2004) but for Gadanki it ranges from -1.78 to -
129 3.25 (without ‘prewhitening/postcoloring’) and -2.39 to -3.13 (with ‘prewhitening/postcoloring’)
130 (during 19-26 March 2004). The higher spectral slope value of -1.78 or -2.39 and -2.90 (without
131 or with ‘prewhitening/postcoloring’) over Gadanki corresponds to the altitude of 60-66 km and
132 the lower values of -3.25 or -3.13 (without or with ‘prewhitening/postcoloring’) is seen at 40-50
133 km. In this scenario, the mesospheric spectral slopes are much lower than the upper stratospheric
134 slopes which is not observed by Guharay and Sekar, (2011). In general, the spectral slope values
135 are below the theoretical saturated slope value of ‘-3’ (except August and October) (Guharay and
136 Sekar, 2011) which is mostly similar to our observation over both the tropical and sub-tropical
137 station of Gadanki (except at few altitudes) and Mt. Abu, respectively (**Table 1**).

138 The VWN spectral slope values on the common operational days (19 March 2004, 20 March 2004
139 and 21 March 2004) over both Gadanki and Mt. Abu are compared in **Figure 5**, **Figure 6** and
140 **Figure 7** respectively to have an insight of the spectral slope variation with height. The slope
141 values are more negative over Gadanki (tropical station) than Mt. Abu (sub-tropical station) for all
142 the altitudes except for 40-50 km where it is same and 60-70 km where slope value over Mt. Abu
143 is lesser than Gadanki (**Figure 5**). On 21 March 2004, the spectral slopes are more negative over
144 Gadanki than Mt. Abu except for the height of 50-60 km and 60-70 km (**Figure 6**). In **Figure 7**,
145 similar features resembling **Figure 5** is observed where the slope values are nearly same at 40-50
146 km and lesser over Mt. Abu than Gadanki at 60-70 km. Guharay et al., (2010) using radiosonde
147 measurements over Gadanki reported the mean spectral slope to be -2.83 at the heights of 18–30
148 km (in the lower stratosphere). Our observations are similar to that of Guharay et al., (2010) except
149 for the fact in some instances more negative slope values in the upper stratosphere (lesser than the

150 theoretical saturated spectral slope value of '-3' is seen). Yu and Yi, (2008) took simultaneous
151 observation from Raman and Rayleigh lidar (covering 4-60 km) over Wuhan, China (30.5° N,
152 114.4° E) and found the spectral slope value to be nearly '-3'. Sato and Yamada, (2004) also
153 reported the saturated slope value over Shigaraki, Japan (35° N, 136° E) around the theoretical
154 value of '-3' and presumed that the probable effect of 'large-scale wind shear' on VWN spectra.
155 Tsuda et al., (1991) observed the spectral slope in the lower stratosphere (18.5-25 km) ranging
156 between -2.9 to -3.2 in winter months while it ranges between -2.2 to -2.4 during the summer over
157 Shigaraki, Japan (35° N, 136° E) using campaign radiosonde soundings. Allen and Vincent, (1995)
158 used radiosonde data over 18 stations in Australia and Antarctica (covering the latitudinal range
159 of 12° S - 68° S and longitudinal range of 78° E - 159° E) and found the mean spectral slope to be
160 ~ -2.5 (lower stratosphere). Over Andenes, Norway (69.3°N), Wu and Widdel, (1989) reported the
161 slope value (logarithmic) of $\sim -3.0 \pm 0.2$ at the altitudes of 80-95 km using horizontal wind
162 measurements.

163 The spectral slope obtained using ground-based Rayleigh lidar observations are compared with
164 space borne SABER satellite measurements. The altitude profile of VWN spectra obtained from
165 SABER temperature is plotted over Gadanki and Mt. Abu (**Figure 8** and **Figure 9**). **Figure 8**
166 displays the VWN spectra on 18 March over Mt. Abu and 19 March over Gadanki respectively.
167 The slopes over Gadanki are more negative than Mt. Abu (similar to **Figures 5-7**) except for the
168 altitude of 40-50 km where the slope value is -6.97 (for Mt. Abu) and -0.09 (for Gadanki). The
169 presence of stratopause at that altitude along with the location of Mt. Abu at about 1.7 km above
170 mean sea-level which jointly may result in such distinct observation. Yan et al., (2019) found the
171 spectral slope ranging from -1.6 to -3.0 whereby mostly the slope values are less negative than the
172 theoretical value of '-3' using Constellation Observing System for Meteorology, Ionosphere and
173 Climate (COSMIC) temperature data (during January 2007 to February 2014) revealing a 'quasi-
174 biennial cycle' at low latitudes, 'annual cycle' at middle and high latitudes, and 'semi-annual
175 cycle' at boreal high latitudes. Further, they also provided evidence that the spectral slope barely
176 has any longitudinal variation and strongly proposed the slope to be of 'universal nature' where
177 the wave propagation processes (for example, 'wave-wave' or 'wave-background flow'
178 interaction) plays a dominant role rather than the GW sources. Zhang et al., (2017a) used 11 years
179 of radiosonde data (1998-2008) over 92 stations of United States (in the Northern Hemisphere)
180 and reported that the 'vertical wind spectrum' is less negative than the 'horizontal wind spectrum'

181 over a wide latitudinal region, with slopes varying between -1.1 to -0.2 in the troposphere and -0.6
182 to 0.1 in the lower stratosphere. Although our observations are not exactly similar to that of Yan
183 et al., (2019) and Zhang et al., (2017a), their results support the shallow VWN spectral slope value
184 of -1.78 or -2.90 (lidar observation) and -0.09 or -1.53 (without or with
185 ‘prewhitening/postcoloring’) (SABER/TIMED observation) over Gadanki in the present study.
186 However, the much negative slope of -6.97 (-4.97) (without or with ‘prewhitening/postcoloring’)
187 observed with SABER/TIMED observations over Mt. Abu at the altitudes of 40-50 km can only
188 be explained by the assumptions that this can be due to the proximity of stratopause which is a
189 ‘transition zone’ between two atmospheric layers presumably interfering with the generation and
190 breaking of waves (at present and it need to be investigated further in future). The ambiguity could
191 not be due to the linear fit of five points because in that case all the other height regions with
192 similar number of data points also would have shown some discrepancy. Since it is only seen at
193 one altitude range over both the tropical and sub-tropical stations, it need to be looked into in detail
194 in future with more simultaneous observations. Ghosh et al., (2018) found the presence of positive
195 slope (at the altitudes of 12-17 km) and concluded that it may be due to the presence of tropopause
196 which may instigate the generation of new GWs. In **Figure 9**, no such discrete slope is observed
197 and it is seen that the slope values are more negative at Gadanki in the altitude range of 40-50 km
198 and 60-70 km and vice-versa. The altitude profile of amplitude and phase of the dominant wave
199 oscillations over Gadanki (during 19-26 March 2004) and Mt. Abu (during 14-21 March 2004) is
200 plotted in **Figure 10**.

201 The amplitudes of atmospheric gravity wave (GW) increases exponentially due to decrease in
202 atmospheric density as they proliferate from the lower to the higher atmosphere. Throughout their
203 propagation, these GWs divulge energy to the surroundings while their ‘power of wavenumber
204 spectra’ and ‘spectral slope’ gets saturated to conserve the overall stability. This resulted in the
205 theory of universal wavenumber spectra with three distinct regions: unsaturated (log-log slope of
206 2), saturated (log-log slope of -3), and turbulence region (log-log slope of -5/3). Therefore, the
207 GWs oscillation and vertical wavenumber spectral slope are related. If continuous temperature or
208 wind measurements are available, it can be utilized to study the dominant GW oscillations present
209 during the observation period. There have been various studies in past (Ghosh and Ramkumar,
210 2014; Gong et al., 2018; Ma et al., 2017) which linked some particular wave oscillations (e.g. 2-
211 day, 5-day, 3-day) with the vertical wave number spectral slope or with certain dynamic events

212 like sudden stratospheric warming (SSW), convection, etc. obtained over that particular station. In
213 the present study, continuous nightly mean temperature observations are available for 8 days
214 (where the lidar is run for 2-3 hours) at both the subtropical and tropical locations. This provides
215 an opportunity to see if planetary waves also can be linked with vertical wavenumber spectral
216 slope or not. Therefore, in the present study, we made an attempt to link the vertical wavenumber
217 spectra with the dominant wave oscillations (for e.g., how the amplitude, phase and vertical
218 wavelength varies) over both stations. It is seen that about 4-day, 2.7-day and 2-day oscillations
219 dominant over both the stations during the observation period. One interesting feature is that the
220 2-day oscillation shows standing nature over Gadanki (tropical station at 375 m elevation
221 surrounded by small terrains) as well as Mt. Abu (sub-tropical station at an elevation of about 1.7
222 km). The vertical wavelength of about 4-day, 2.7-day and 2-day oscillation calculated from the
223 altitude profile of phase is tabulated in **Table 1**. It is to be noted that the 2-day wave have a single
224 vertical wavelength value throughout the altitude profile owing to its standing nature. The vertical
225 wavelength (km) over Gadanki (Mt. Abu): for 4-day oscillation ranges from 3.52-25.54 (8.37-
226 43.64) km, for 2.7 day oscillation 4.89-17.81 (3.87-16.41) km, and for 2-day oscillation it is 22.19
227 (29.82) km.

228 **4. Summary and conclusions**

229 The present study describes in detail the VWN spectral slope characteristics over a tropical station
230 of Gadanki and sub-tropical station of Mt. Abu. It is observed that the spectral slopes over Gadanki
231 are more negative in most of the altitudes with some exceptions. This could be due to the fact that
232 over Gadanki, the convectively generated gravity waves play a major role whereas in case of Mt.
233 Abu, mountain generated gravity waves affects the atmosphere in majority. Although more
234 negative slope value than the theoretical saturated slope of '-3' is observed over Gadanki, they are
235 much less than those reported by Ghosh et al., (2018) in the order of ~ -6 , ~ -9 and ~ -12 . Moreover,
236 in the present case no positive spectral slope is observed indicating wave generation (Ghosh et al.,
237 2018). The spectral slopes observed in the present study depict similar spectral slope nature over
238 Gadanki as reported by earlier studies using radiosonde (Guharay et al., 2010) and Rayleigh lidar
239 (Guharay and Sekar, 2011) over Gadanki. To investigate convection generated gravity wave effect
240 on VWN spectra, further studies need to be done with large datasets in future and other instruments
241 (as Rayleigh lidar cannot be operated during cloudy and rainy nights). Our observations even
242 match with the earlier studies made with COSMIC temperature, radiosonde and radio occultation

243 measurements in the mid and high latitudes (Zhang et al., 2017a; Tsuda and Hocke, 2002; Yan et
244 al., 2019). Since lidar observations are made simultaneously over both subtropical and tropical
245 stations during similar period, it gave an opportunity to study planetary wave oscillations over both
246 the locations and how their characteristics vary over each locations. Similar wave oscillations are
247 observed over both the stations with periodicities of about 4-day, 2.7-day and 2-day with
248 comparable vertical wavelength. One interesting feature is that the 2-day wave shows standing
249 nature over both the tropical and sub-tropical locations.

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356 **Figure Captions**

357 **Figure 1.** Altitude profile (33-70 km) of vertical wavenumber spectra of Rayleigh lidar
358 temperature during 14-17 March 2004 over Mt. Abu (24.5° N, 72.7° E) [14 March: a-d; 15 March:
359 e-h; 16 March: i-l; 17 March: m-p]. The slope lines and values are calculated considering both
360 without and with ‘prewhitening/postcoloring’ technique where the first slope region (Slope 1;
361 denoted with magenta color) indicates slope value without ‘prewhitening/postcoloring’ and the
362 second slope region (Slope 2; denoted with blue color) indicates slope value with
363 ‘prewhitening/postcoloring’.

364 **Figure 2.** Same as Figure 1 except for during 18-21 March 2004 [18 March: a-d; 19 March: e-h;
365 20 March: i-l; 21 March: m-p].

366 **Figure 3.** Altitude profile (30-66 km) of vertical wavenumber spectra of Rayleigh lidar
367 temperature during 19-22 March 2004 over Gadanki (13.5° N, 79.2° E) [19 March: a-d; 20 March:
368 e-h; 21 March: i-l; 22 March: m-p]. The slope lines and values are calculated considering both
369 without and with ‘prewhitening/postcoloring’ technique where the first slope region (Slope 1;
370 denoted with magenta color) indicates slope value without ‘prewhitening/postcoloring’ and the
371 second slope region (Slope 2; denoted with blue color) indicates slope value with
372 ‘prewhitening/postcoloring’.

373 **Figure 4.** Same as Figure 3 except for during 23-26 March 2004 [23 March: a-d; 24 March: e-h;
374 25 March: i-l; 26 March: m-p].

375 **Figure 5.** Comparison of vertical wavenumber spectra of height profile of Rayleigh lidar
376 temperature over Mt. Abu [33-40 km (d); 40-50 km (c); 50-60 km (b); 60-70 km (a)] and Gadanki
377 [30-40 km (h); 40-50 km (g); 50-60 km (f); 60-66 km (e)] on 19 March 2004 [Slope 1 (denoted
378 with magenta color) indicates slope value without ‘prewhitening/postcoloring’ and Slope 2
379 (denoted with blue color) indicates slope value with ‘prewhitening/postcoloring’].

380 **Figure 6.** Same as Figure 5 except for on 20 March 2004.

381 **Figure 7.** Same as Figure 5 except for on 21 March 2004.

382 **Figure 8.** Comparison of altitude profile (30-70 km) of vertical wavenumber spectra of SABER
383 temperature on 18 March 2004 over Mt. Abu [30-40 km (d); 40-50 km (c); 50-60 km (b); 60-70
384 km (a)] and 19 March 2004 over Gadanki [30-40 km (h); 40-50 km (g); 50-60 km (f); 60-70 km
385 (e)]. Slope value without ‘prewhitening/postcoloring’ is represented as Slope 1 (with magenta
386 color) and with ‘prewhitening/postcoloring’ as Slope 2 (with blue color).

387 **Figure 9.** Same as Figure 8 except for 21 March 2004 over Mt. Abu and 20 March 2004 over
388 Gadanki.

389 **Figure 10.** Height profile of amplitude and phase of wave oscillations in Rayleigh lidar
390 temperature over Gadanki (during 19-26 March 2004; top panel) and Mt. Abu (during 14-21 March
391 2004; lower panel). The dominant oscillations observed are shown in different colors (blue color:
392 4-day oscillation, green color: 2.7-day oscillation and wine color: 2-day oscillation).

393

394 **Table Captions**

395 **Table 1.** Vertical wavenumber (VWN) spectral slope (approximate values) using lidar
396 observations without Pre-whitening and Post-Coloring (PWPC)] and with PWPC over: A) Mt.
397 Abu (24.5° N, 72.7° E), and B) Gadanki (13.5° N, 79.2° E). The spectral slopes in bracket indicates
398 VWN spectral slope with PWPC.

399 **Table 2.** Co-efficient of determination (R^2) of vertical wavenumber spectral slopes using lidar
400 observation [with and without Pre-whitening and Post-Coloring (PWPC)] over: A) Mount Abu
401 (24.5° N, 72.7° E), and B) Gadanki (13.5° N, 79.2° E). The values denoted with ‘±’ sign are standard
402 deviation of R^2 .

403 **Table 3.** (i) Vertical wavenumber (VWN) spectral slope (approximate values) without and with
404 Pre-whitening and Post-Coloring (PWPC)], (ii) Co-efficient of determination (R^2) obtained using
405 SABER/TIMED observations without and with Pre-whitening and Post-Coloring (PWPC)] over
406 Mt. Abu (24.5° N, 72.7° E) on 18 and 21 March 2004 and Gadanki (13.5° N, 79.2° E) on 19 and
407 20 March 2004. The spectral slopes in bracket indicates VWN spectral slope and Co-efficient of
408 determination (R^2) with PWPC.

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411 **Table 1.** Vertical wavenumber (VWN) spectral slope (approximate values) using lidar
 412 observations without Pre-whitening and Post-Coloring (PWPC)] and with PWPC over: A) Mt.
 413 Abu (24.5° N, 72.7° E), and B) Gadanki (13.5° N, 79.2° E). The spectral slopes in bracket indicates
 414 VWN spectral slope with PWPC.

A) Mount Abu (24.5° N, 72.7° E)								
Altitude Range (km)	14 March	15 March	16 March	17 March	18 March	19 March	20 March	21 March
33-40	-2.63 (-2.81)	-2.57 (-2.78)	-2.68 (-2.84)	-2.64 (-2.81)	-2.66 (-2.82)	-2.73 (-2.86)	-2.61 (-2.80)	-2.68 (-2.83)
40-50	-2.82 (-2.91)	-2.67 (-2.83)	-2.77 (-2.88)	-2.80 (-2.90)	-2.50 (-2.75)	-2.94 (-2.97)	-2.86 (-2.93)	-2.93 (-2.96)
50-60	-2.46 (-2.73)	-2.43 (-2.71)	-2.88 (-2.94)	-2.14 (-2.57)	-2.79 (-2.89)	-2.42 (-2.71)	-2.66 (-2.83)	-2.52 (-2.76)
60-70	-2.72 (-2.86)	-2.65 (-2.82)	-2.72 (-2.83)	-2.58 (-2.79)	-2.75 (-2.87)	-2.68 (-2.84)	-2.75 (-2.81)	-2.79 (-2.89)
B) Gadanki (13.5° N, 79.2° E)								
Altitude Range (km)	19 March	20 March	21 March	22 March	23 March	24 March	25 March	26 March
30-40	-2.94 (-2.97)	-2.71 (-2.86)	-2.83 (-2.91)	-3.01 (-3.00)	-2.71 (-2.86)	-2.86 (-2.93)	-2.69 (-2.84)	-2.50 (-2.75)
40-50	-2.94 (-2.97)	-3.01 (-3.01)	-2.99 (-2.99)	-3.25 (-3.13)	-2.94 (-2.97)	-2.82 (-2.90)	-3.17 (-3.09)	-2.78 (-2.89)
50-60	-2.86 (-2.93)	-2.53 (-2.76)	-2.93 (-2.96)	-2.68 (-2.84)	-2.56 (-2.78)	-2.40 (-2.70)	-2.47 (-2.73)	-2.74 (-2.87)
60-66	-2.31 (-2.65)	-2.40 (-2.70)	-2.31 (-2.58)	-2.30 (-2.65)	-1.78 (-2.39)	-2.70 (-2.85)	-1.78 (-2.90)	-2.36 (-2.68)

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422 **Table 2.** Co-efficient of determination (R^2) of vertical wavenumber spectral slopes using lidar
 423 observation [with and without Pre-whitening and Post-Coloring (PWPC)] over: A) Mount Abu
 424 (24.5° N, 72.7° E), and B) Gadanki (13.5° N, 79.2° E). The values denoted with '±' sign are standard
 425 deviation of R^2 .

A) Mount Abu (24.5° N, 72.7° E)								
<i>i) Co-efficient of determination (R^2) [without Pre-whitening and Post-Coloring (PWPC)]</i>								
Altitude Range (km)	14 March	15 March	16 March	17 March	18 March	19 March	20 March	21 March
33-40	0.995 ± 0.01	0.998 ± 0.01	0.987 ± 0.01	0.994 ± 0.02	0.995 ± 0.01	0.992 ± 0.01	0.996 ± 0.01	0.996 ± 0.03
40-50	0.982 ± 0.01	0.986 ± 0.01	0.975 ± 0.01	0.989 ± 0.02	0.976 ± 0.01	0.984 ± 0.01	0.968 ± 0.01	0.989 ± 0.03
50-60	0.984 ± 0.01	0.982 ± 0.01	0.972 ± 0.01	0.952 ± 0.02	0.967 ± 0.01	0.980 ± 0.01	0.988 ± 0.01	0.935 ± 0.03
60-70	0.967 ± 0.01	0.979 ± 0.01	0.967 ± 0.01	0.987 ± 0.02	0.974 ± 0.01	0.996 ± 0.01	0.974 ± 0.01	0.998 ± 0.03
<i>ii) Co-efficient of determination (R^2) [with Pre-whitening and Post-Coloring (PWPC)]</i>								
Altitude Range (km)	14 March	15 March	16 March	17 March	18 March	19 March	20 March	21 March
33-40	0.999 ± 0.00	1.00 ± 0.00	0.997 ± 0.00	0.999 ± 0.00	0.999 ± 0.00	0.998 ± 0.00	0.999 ± 0.00	0.999 ± 0.01
40-50	0.996 ± 0.00	0.997 ± 0.00	0.994 ± 0.00	0.997 ± 0.00	0.995 ± 0.00	0.996 ± 0.00	0.992 ± 0.00	0.997 ± 0.01
50-60	0.997 ± 0.00	0.996 ± 0.00	0.993 ± 0.00	0.991 ± 0.00	0.992 ± 0.00	0.996 ± 0.00	0.997 ± 0.00	0.986 ± 0.01
60-70	0.992 ± 0.00	0.995 ± 0.00	0.991 ± 0.00	0.997 ± 0.00	0.994 ± 0.00	0.999 ± 0.00	0.996 ± 0.00	0.999 ± 0.01
B) Gadanki (13.5° N, 79.2° E)								
<i>i) Co-efficient of determination (R^2) [without Pre-whitening and Post-Coloring (PWPC)]</i>								
Altitude Range (km)	19 March	20 March	21 March	22 March	23 March	24 March	25 March	26 March
30-40	0.904 ± 0.03	0.865 ± 0.09	0.973 ± 0.09	0.956 ± 0.13	0.973 ± 0.18	0.976 ± 0.14	0.976 ± 0.16	0.888 ± 0.04
40-50	0.948 ± 0.03	0.957 ± 0.09	0.778 ± 0.09	0.949 ± 0.13	0.931 ± 0.18	0.783 ± 0.14	0.809 ± 0.16	0.912 ± 0.04
50-60	0.882 ± 0.03	0.751 ± 0.09	0.794 ± 0.09	0.672 ± 0.13	0.882 ± 0.18	0.817 ± 0.14	0.812 ± 0.16	0.898 ± 0.04

60-66	0.905 ± 0.03	0.802 ± 0.09	0.905 ± 0.09	0.812 ± 0.13	0.582 ± 0.18	0.632 ± 0.14	0.582 ± 0.16	0.825 ± 0.04
<i>ii) Co-efficient of determination (R²) [with Pre-whitening and Post-Coloring (PWPC)]</i>								
Altitude Range (km)	19 March	20 March	21 March	22 March	23 March	24 March	25 March	26 March
30-40	0.975 ± 0.01	0.966 ± 0.02	0.994 ± 0.03	0.989 ± 0.04	0.994 ± 0.04	0.994 ± 0.05	0.994 ± 0.02	0.975 ± 0.01
40-50	0.987 ± 0.01	0.989 ± 0.02	0.934 ± 0.03	0.986 ± 0.04	0.982 ± 0.04	0.939 ± 0.05	0.941 ± 0.02	0.978± 0.01
50-60	0.969 ± 0.01	0.935 ± 0.02	0.940 ± 0.03	0.902 ± 0.04	0.972 ± 0.04	0.957 ± 0.05	0.955 ± 0.02	0.975 ± 0.01
60-66	0.980 ± 0.01	0.953 ± 0.02	0.938 ± 0.03	0.958 ± 0.04	0.909 ± 0.04	0.884 ± 0.05	0.976 ± 0.02	0.960 ± 0.01

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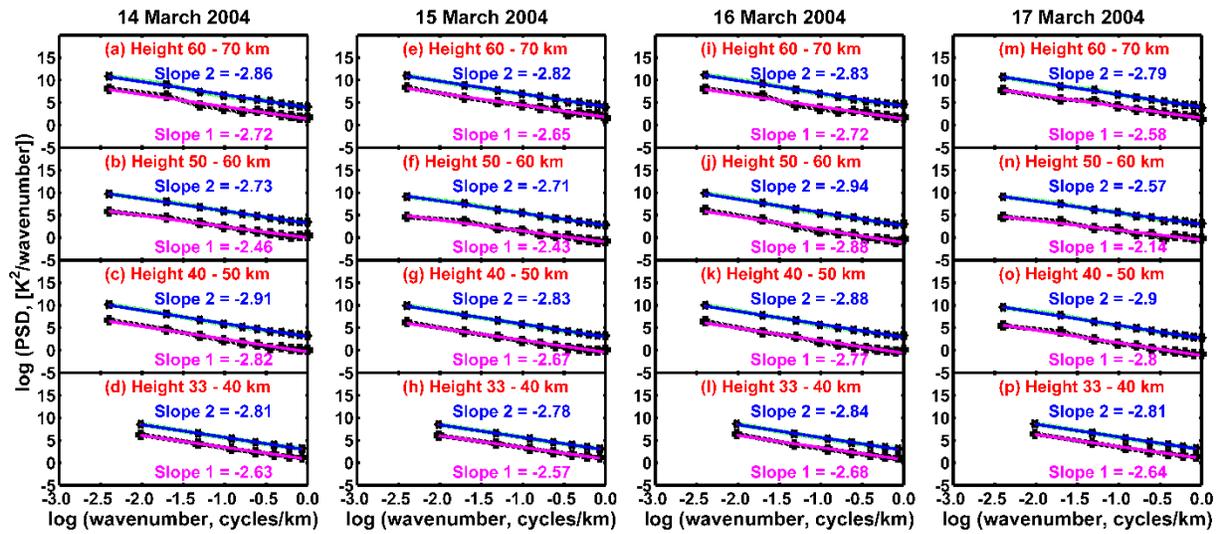
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441 **Table 3.** (i) Vertical wavenumber (VWN) spectral slope (approximate values) without and with
 442 Pre-whitening and Post-Coloring (PWPC)], (ii) Co-efficient of determination (R^2) obtained using
 443 SABER/TIMED observations without and with Pre-whitening and Post-Coloring (PWPC)] over
 444 Mt. Abu (24.5° N, 72.7° E) on 18 and 21 March 2004 and Gadanki (13.5° N, 79.2° E) on 19 and
 445 20 March 2004. The spectral slopes in bracket indicates VWN spectral slope and Co-efficient of
 446 determination (R^2) with PWPC.

<i>i) Vertical wavenumber (VWN) spectral slope (approximate values)</i>				
Altitude Range (km)	18 March (Mt. Abu)	19 March (Gadanki)	21 March (Mt. Abu)	20 March (Gadanki)
30-40	-2.61 (-2.79)	-2.70 (-2.84)	-2.49 (-2.73)	-2.30 (-2.64)
40-50	-6.97 (-4.97)	-0.09 (-1.53)	-2.60 (-2.78)	-2.75 (-2.86)
50-60	-2.34 (-2.66)	-2.97 (-2.97)	-2.82 (-2.90)	-2.36 (-2.67)
60-70	-2.43 (-2.70)	-2.63 (-2.80)	-2.19 (-2.58)	-2.71 (-2.84)
<i>ii) Co-efficient of determination (R^2)</i>				
Altitude Range (km)	18 March (Mt. Abu)	19 March (Gadanki)	21 March (Mt. Abu)	20 March (Gadanki)
30-40	0.998 ± 0.04 (0.999 ± 0.02)	0.987 ± 0.49 (0.997 ± 0.17)	0.992 ± 0.01 (0.998 ± 0.00)	0.993 ± 0.00 (0.998 ± 0.00)
40-50	0.907 ± 0.04 (0.952 ± 0.02)	0.175 ± 0.49 (0.656 ± 0.17)	0.989 ± 0.01 (0.997 ± 0.00)	0.992 ± 0.00 (0.998 ± 0.00)
50-60	0.949 ± 0.04 (0.989 ± 0.02)	0.974 ± 0.49 (0.993 ± 0.17)	0.978 ± 0.01 (0.994 ± 0.00)	0.983 ± 0.00 (0.996 ± 0.00)
60-70	0.985 ± 0.04 (0.997 ± 0.02)	0.998 ± 0.49 (0.999 ± 0.17)	0.996 ± 0.01 (0.999 ± 0.00)	0.993 ± 0.00 (0.998 ± 0.00)

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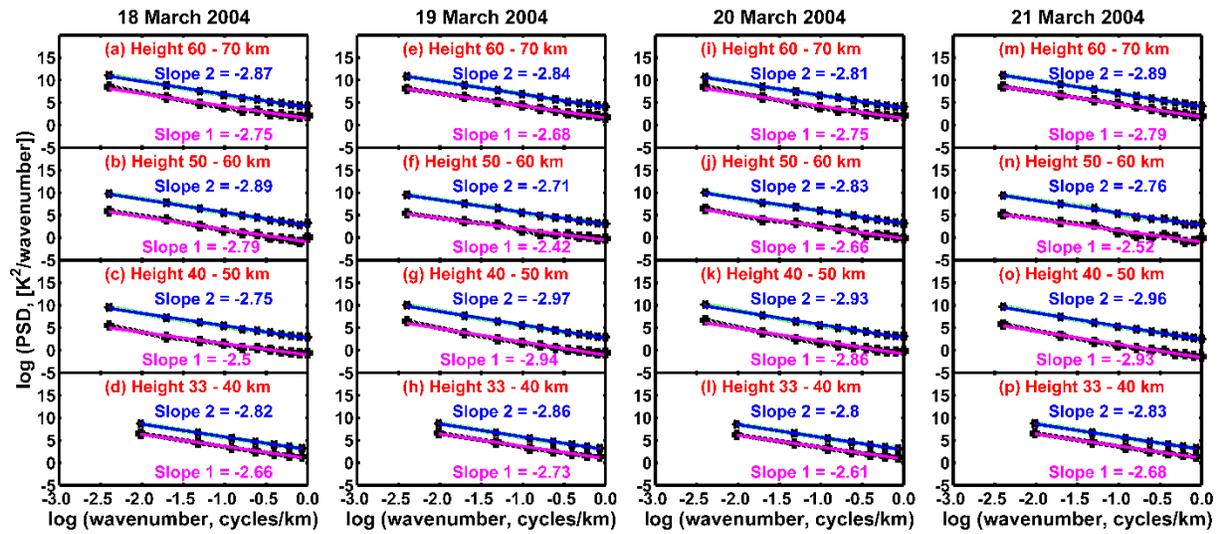
449 **Figure 1.** Altitude profile (33-70 km) of vertical wavenumber spectra of Rayleigh lidar
 450 temperature during 14-17 March 2004 over Mt. Abu (24.5° N, 72.7° E) [14 March: a-d; 15 March:
 451 e-h; 16 March: i-l; 17 March: m-p]. The slope lines and values are calculated considering both
 452 without and with ‘prewhitening/postcoloring’ technique where the first slope region (Slope 1;
 453 denoted with magenta color) indicates slope value without ‘prewhitening/postcoloring’ and the
 454 second slope region (Slope 2; denoted with blue color) indicates slope value with
 455 ‘prewhitening/postcoloring’.

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461 **Figure 2.** Same as Figure 1 except for during 18-21 March 2004 [18 March: a-d; 19 March: e-h;
 462 20 March: i-l; 21 March: m-p].

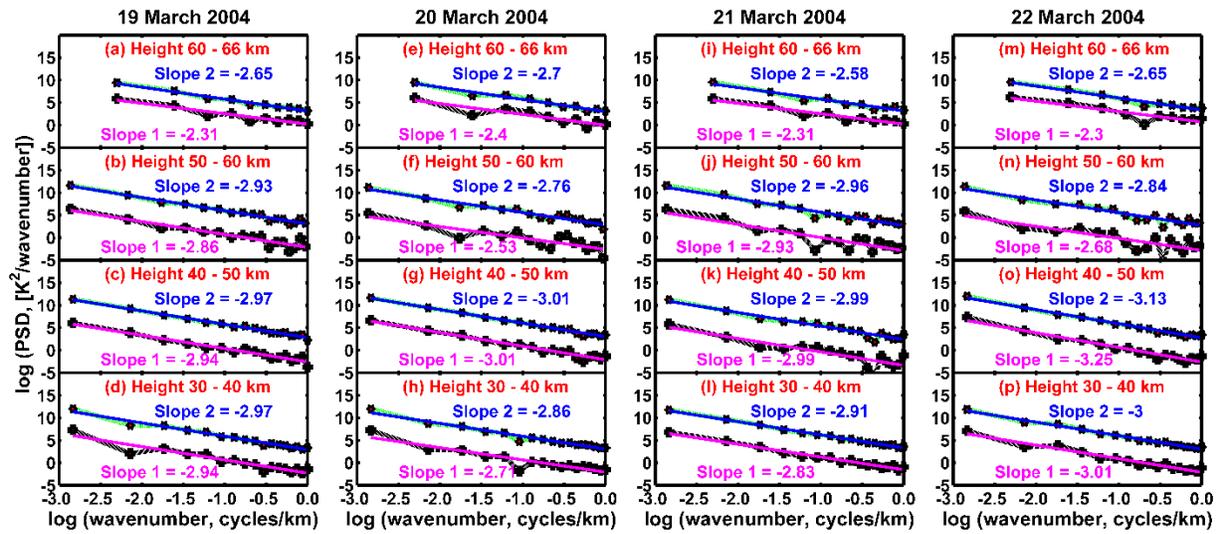
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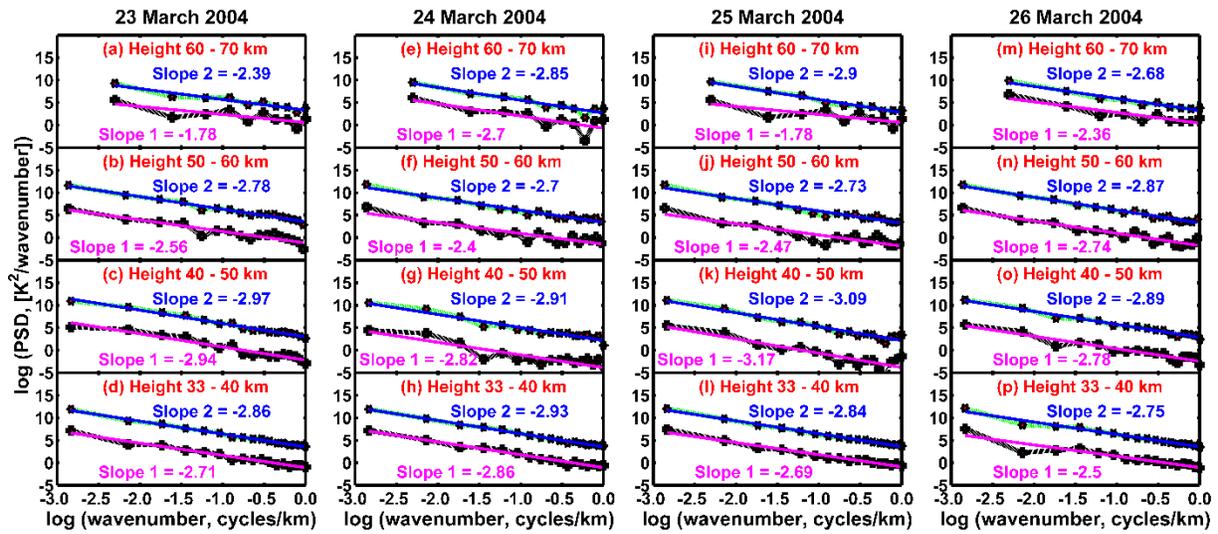


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469 **Figure 3.** Altitude profile (30-66 km) of vertical wavenumber spectra of Rayleigh lidar
 470 temperature during 19-22 March 2004 over Gadanki (13.5° N, 79.2° E). [19 March: a-d; 20 March:
 471 e-h; 21 March: i-l; 22 March: m-p]. The slope lines and values are calculated considering both
 472 without and with ‘prewhitening/postcoloring’ technique where the first slope region (Slope 1;
 473 denoted with magenta color) indicates slope value without ‘prewhitening/postcoloring’ and the
 474 second slope region (Slope 2; denoted with blue color) indicates slope value with
 475 ‘prewhitening/postcoloring’.

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479 **Figure 4.** Same as Figure 3 except for during 23-26 March 2004 [23 March: a-d; 24 March: e-h;
 480 25 March: i-l; 26 March: m-p].

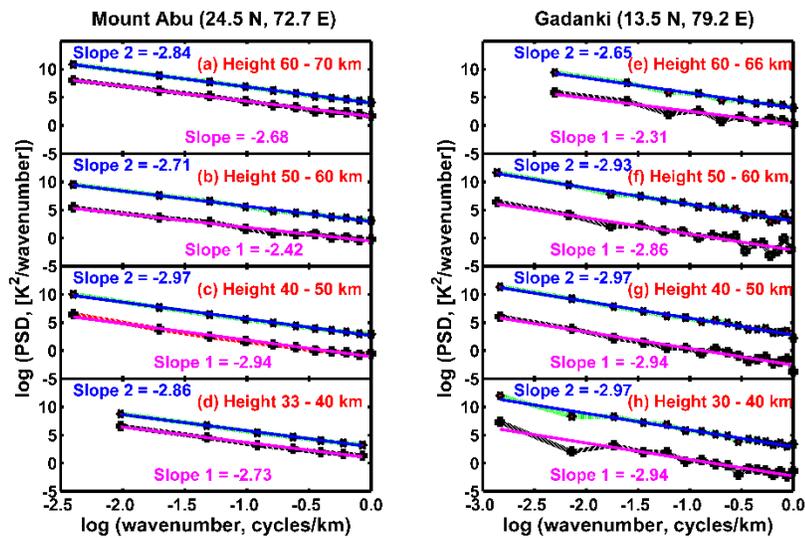
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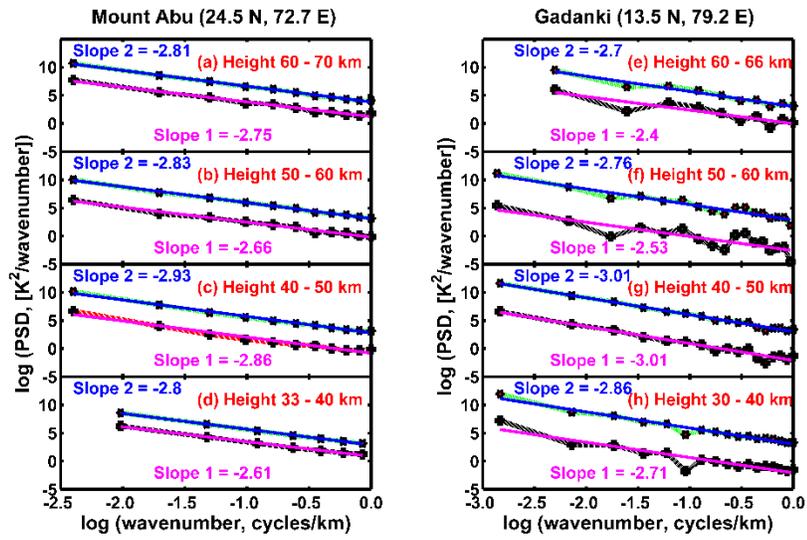
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487 **Figure 5.** Comparison of vertical wavenumber spectra of height profile of Rayleigh lidar
 488 temperature over Mt. Abu [33-40 km (d); 40-50 km (c); 50-60 km (b); 60-70 km (a)] and Gadanki
 489 [30-40 km (h); 40-50 km (g); 50-60 km (f); 60-66 km (e)] on 19 March 2004 [Slope 1 (denoted
 490 with magenta color) indicates slope value without ‘prewhitening/postcoloring’ and Slope 2
 491 (denoted with blue color) indicates slope value with ‘prewhitening/postcoloring’].

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496 **Figure 6.** Same as Figure 5 except for on 20 March 2004.

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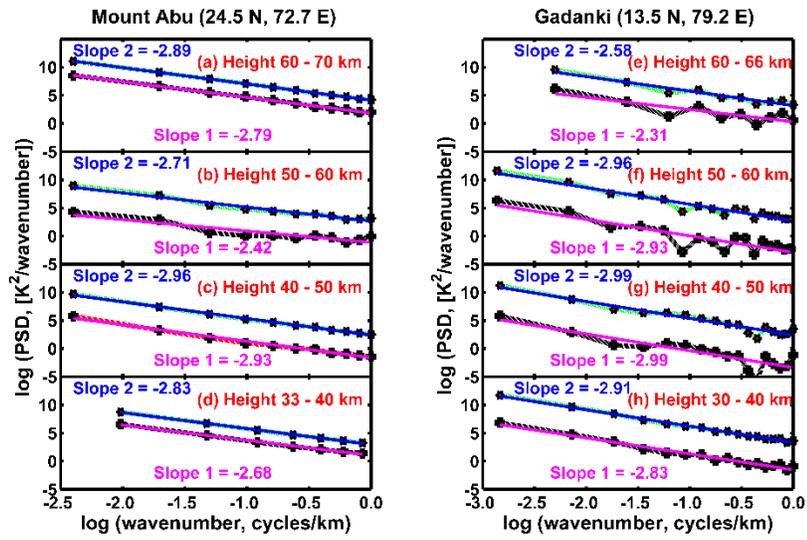
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504 **Figure 7.** Same as Figure 5 except for on 21 March 2004.

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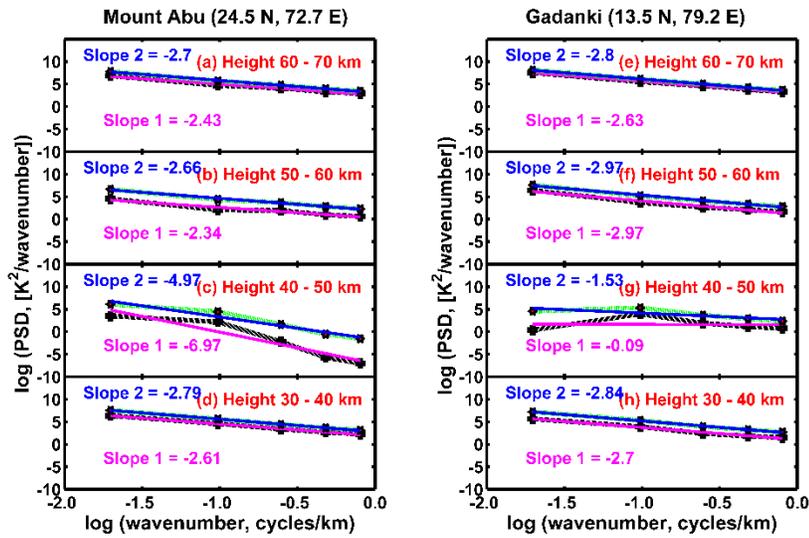
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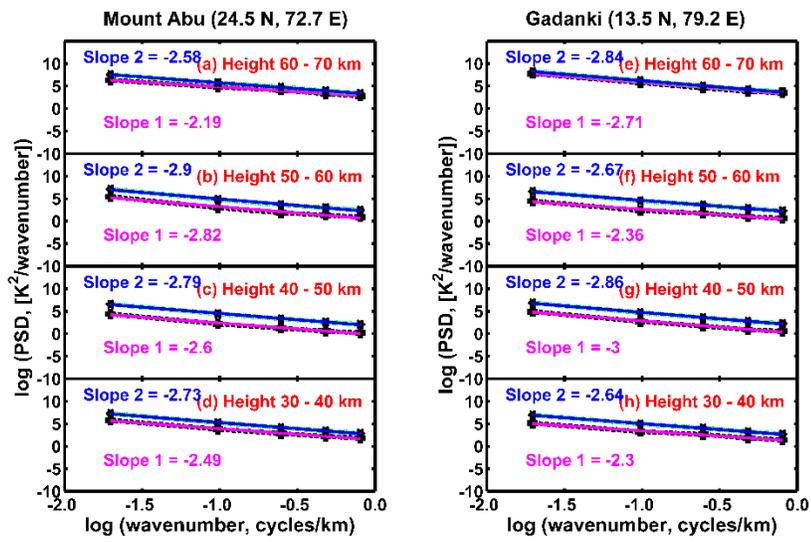
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512 **Figure 8.** Comparison of altitude profile (30-70 km) of vertical wavenumber spectra of SABER
 513 temperature on 18 March 2004 over Mt. Abu [30-40 km (d); 40-50 km (c); 50-60 km (b); 60-70
 514 km (a)] and 19 March 2004 over Gadanki [30-40 km (h); 40-50 km (g); 50-60 km (f); 60-70 km
 515 (e)]. Slope value without ‘prewhitening/postcoloring’ is represented as Slope 1 (with magenta
 516 color) and with ‘prewhitening/postcoloring’ as Slope 2 (with blue color).

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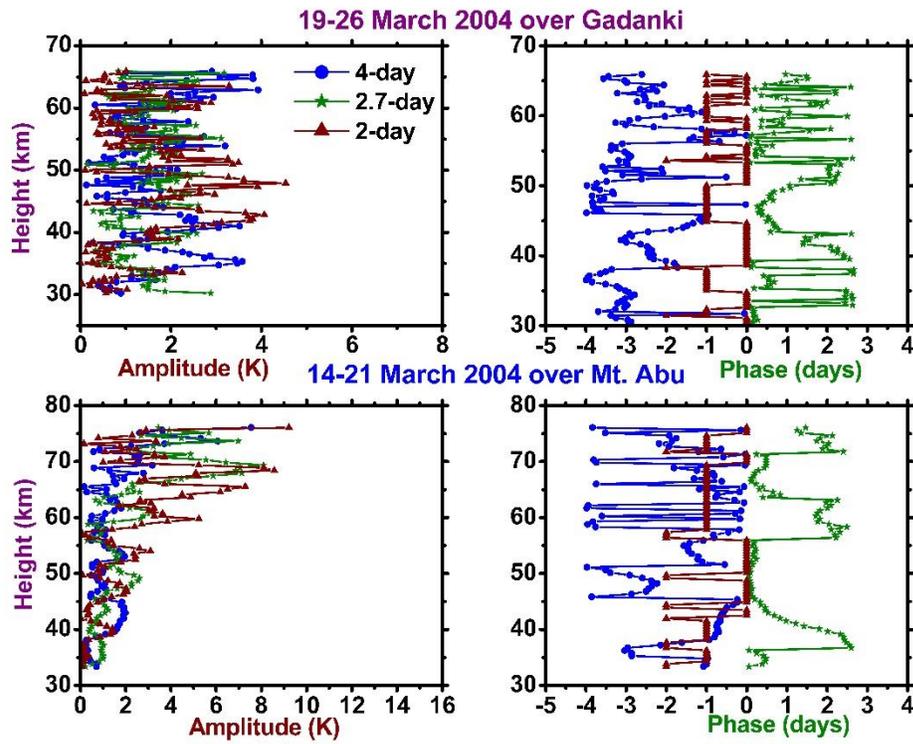
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521 **Figure 9.** Same as Figure 8 except for 21 March 2004 over Mt. Abu and 20 March 2004 over
 522 Gadanki.

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527 **Figure 10.** Height profile of amplitude and phase of wave oscillations in Rayleigh lidar
 528 temperature over Gadanki (during 19-26 March 2004; top panel) and Mt. Abu (during 14-21 March
 529 2004; lower panel). The dominant oscillations observed are shown in different colors (blue color:
 530 4-day oscillation, green color: 2.7-day oscillation and wine color: 2-day oscillation).