

Theory and simulations of relativistic particle motions in a magnetosonic shock wave

SHUNSUKE USAMI¹ and Y. OHSAWA²

¹Computer and Information Network Center, National Institute for Fusion Science,
322-6 Oroshi-cho, Toki, Gifu 509-5292, Japan

²Department of Physics, Nagoya University, Nagoya 464-8602, Japan

(Received 15 August 2005 and accepted 20 December 2005)

Abstract. The motions of relativistic particles in a magnetosonic shock wave propagating obliquely to an external magnetic field are studied. In the zeroth-order theory, particles continue to move nearly parallel to the external magnetic field in the shock transition region, when the shock speed is close to $c \cos \theta$, where c is the speed of light and θ is the propagation angle. Perturbations to this zeroth-order motion are also analyzed for positrons and ions. The perturbation frequency of positrons is $\omega \sim \Omega_{p0} \gamma^{-1}$ and that of ions is $\omega \sim \Omega_{i0} \gamma^{-1/2}$, where Ω_{p0} and Ω_{i0} are the non-relativistic gyrofrequencies of positrons and of ions, respectively, and γ is the Lorentz factor. These theoretical predictions are confirmed with numerical simulations.

1. Introduction

Magnetosonic shock waves can accelerate thermal hydrogen ions, heavy ions, and electrons with various non-stochastic mechanisms [1]. Recently, two acceleration mechanisms of non-thermal, relativistic particles in magnetosonic shock waves have been studied with theory and simulations. One mechanism is associated with large-radius gyromotions; particles absorb energy from the electric field perpendicular to the magnetic field \mathbf{B} [2]. The other is the acceleration parallel to \mathbf{B} [3, 4]. Here, we investigate particle motions in the latter. In Sec. 2, we describe the zeroth-order and perturbation theories for relativistic particles. In Sec. 3, we verify them with simulations. Section 4 gives a summary of our work.

2. Theory

2.1. Zeroth-order theory

We analyze the motions of relativistic particles in a magnetosonic shock wave propagating in the x direction with a speed v_{sh} in an external magnetic field $\mathbf{B}_0 = B_0(\cos \theta, 0, \sin \theta)$. If the particle speed v is very close to the speed of light c , a slight change in the particle speed can lead to a great change in the Lorentz factor γ . Therefore, ignoring $\gamma d\mathbf{v}/dt$ compared with $\mathbf{v}d\gamma/dt$, we obtain the zeroth-order equation of motion for a particle with a mass m_j and a charge q_j (j denotes particle species, $j = p$ or i):

$$m_j \frac{d\gamma_0}{dt} \mathbf{v}_0 = q_j \left(\mathbf{E} + \frac{\mathbf{v}_0}{c} \times \mathbf{B} \right), \quad (2.1)$$

where the subscript 0 refers to the zeroth-order quantities.

We consider particles moving with the wave, i.e. $v_{x0} = v_{\text{sh}}$. When $v_{\text{sh}} \approx c \cos \theta$, we find

$$\frac{v_{z0}}{v_{\text{sh}}} \approx \frac{B_{z0}}{B_{x0}} \quad (2.2)$$

from (2.1). It is also shown that $|v_{y0}|$ is much smaller than v_{x0} and v_{z0} ; thus, the particles move nearly parallel to \mathbf{B}_0 . Moreover, we obtain the time rate of change of γ_0 as

$$\frac{d\gamma_0}{dt} = \frac{q_j B_{x0} (\mathbf{E} \cdot \mathbf{B}_0)}{m_j v_{\text{sh}} (\mathbf{B} \cdot \mathbf{B}_0)}. \quad (2.3)$$

If the particle position in the wave does not change, γ_0 continues to grow linearly with time [3].

2.2. Perturbation theory

The zeroth-order theory is applicable to either positrons or ions. We do, however, need to treat positron and ion perturbations separately [4]. We assume that $d\mathbf{v}_1/dt \sim \gamma_0^{-1} \Omega_{p0} \mathbf{v}_1$ for positrons and $d\mathbf{v}_1/dt \sim \gamma_0^{-1/2} \Omega_{i0} \mathbf{v}_1$ for ions, where Ω_{p0} and Ω_{i0} are the non-relativistic positron and ion gyrofrequencies, respectively, and the subscript 1 refers to perturbed quantities. We then expand the exact equation of motion. After some algebra, the perturbation frequency ω of the positrons is obtained as

$$\omega^2 = \left(\frac{e}{m_p c} \right)^2 \frac{(\gamma_0^2/c^2)(\mathbf{B}_0 \cdot \mathbf{v}_0)^2 + B^2}{\gamma_0^4}, \quad (2.4)$$

which is obviously positive, while that of the ions is given as

$$\omega^2 = -\frac{q_i}{m_i \gamma_0 \gamma_{\text{sh}}} \left(\frac{dE_x}{d\xi_0} + \frac{v_{y0}}{c} \frac{dB_z}{d\xi_0} - \frac{v_{z0}}{c} \frac{dB_y}{d\xi_0} \right), \quad (2.5)$$

where $\gamma_{\text{sh}} = [1 - (v_{\text{sh}}/c)^2]^{-1/2}$, $\xi = x - v_{\text{sh}}t$, and ξ_0 is the center position of the perturbation; $d/d\xi_0$ designates the derivative at $\xi = \xi_0$. The ion perturbation is stable when $\omega^2 > 0$.

3. Numerical studies

We numerically investigate the motions of relativistic particles. For positrons, we use a one-dimensional, relativistic, electromagnetic particle simulation code [3]. As in the theory, waves propagate in the x direction in an external magnetic field \mathbf{B}_0 . The field strength is $|\Omega_{e0}|/\omega_{pe} = 3$, where $|\Omega_{e0}|$ and ω_{pe} are the electron gyro and plasma frequencies, respectively. The propagation angle is taken to be $\theta = 42^\circ$. The propagation speed of a shock wave studied here is observed to be $v_{\text{sh}} = 2.4v_A$, where v_A is the Alfvén speed. It has a typical shock profile with the width of the transition region of the order of the ion inertial length as shown, for instance, in [3]. Figure 1 shows the time variation of γ of a positron accelerated by a magnetosonic shock wave in an electron–positron–ion plasma with a positron-to-electron density ratio of 0.02 [3]. The energy increases up to $\gamma \sim 600$. We did, however, find oscillations in γ . Figure 2 displays the oscillation frequency ω as a function of γ_0 [4]. The data points represent simulation results, while the solid line shows the theoretical curve given by (2.4). The simulation results are explained by the theory.

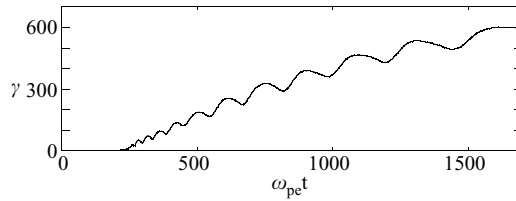


Figure 1. Time variation of γ of an accelerated positron. Here, ω_{pe} is the electron plasma frequency.

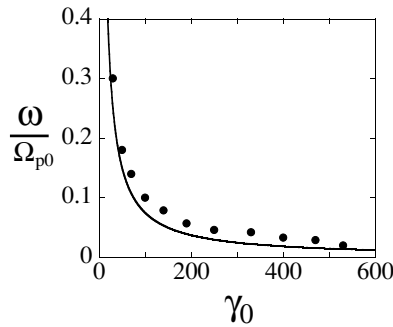


Figure 2. Positron perturbation frequency ω versus γ_0 .

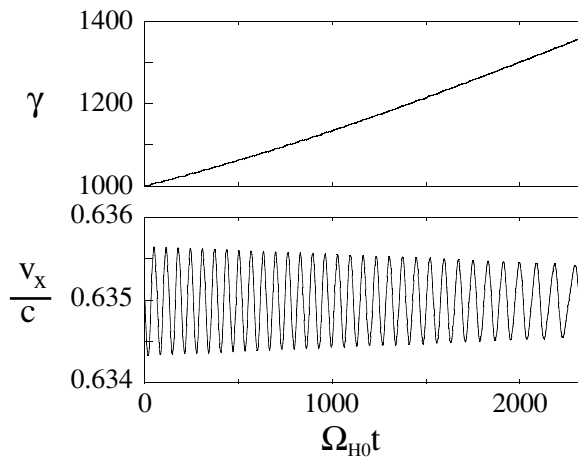


Figure 3. Time variations of γ and v_x of an accelerated hydrogen ion. Here, Ω_{H0} is the non-relativistic hydrogen gyrofrequency.

For ions, we calculate test particle orbits; we first obtain the electromagnetic fields in a shock wave from a particle simulation and then follow particle motions in the fields, assuming stationary wave propagation. Here, plasma parameters are $|\Omega_{e0}|/\omega_{pe} = 1.5$, $\theta = 50^\circ$, and $v_{sh} = 3.2v_A$. The initial velocities of the test particles are given by the zeroth-order theory. In Fig. 3 we show the time variations of γ and v_x of an ion with an initial energy $\gamma = 1000$. The energy increases from $\gamma = 1000$

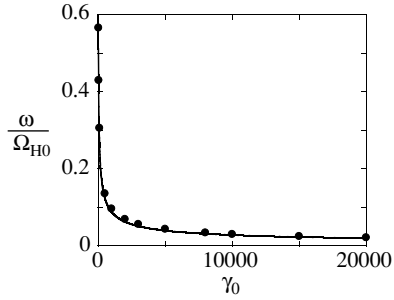


Figure 4. Ion perturbation frequency ω versus γ_0 .

to approximately 1400. Also, we find small-amplitude oscillations in v_x . In Fig. 4, we display the perturbation frequency ω as a function of γ_0 . The simulation values fit well to the theoretical curve obtained from (2.5)

4. Summary

The motions of relativistic particles in a magnetosonic shock wave have been studied. In the zeroth-order theory, where the relation $\gamma|d\mathbf{v}/dt| \ll v d\gamma/dt$ is assumed, particles are accelerated almost parallel to the external magnetic field, when $v_{sh} \approx c \cos \theta$. This is applicable to either positrons or ions. Perturbation theories for positrons and ions have been separately investigated. The perturbation frequency of positrons is $\omega \sim \Omega_{p0} \gamma_0^{-1}$ and that of ions is $\omega \sim \Omega_{i0} \gamma_0^{-1/2}$. The zeroth- and first-order theories have been verified with numerical simulations.

References

- [1] Ohsawa, Y. 2004 Nonstochastic particle acceleration in collisionless shock waves. *Physica Scripta* **T107**, 32–35.
- [2] Usami, S. and Ohsawa, Y. 2004 Evolution of relativistic ions incessantly accelerated by an oblique shock wave. *Phys. Plasmas* **11**, 918–925.
- [3] Hasegawa, H., Usami, S. and Ohsawa, Y. 2003 Positron acceleration to ultrarelativistic energies by a shock wave in a magnetized electron–positron–ion plasma. *Phys. Plasmas* **10**, 3455–3458.
- [4] Usami, S. and Ohsawa, Y. 2004 Motions of ultrarelativistic particles accelerated in an oblique plasma wave. *Phys. Plasmas* **11**, 3203–3211.